

The American Association of Variable Star Observers

AAVSO

Annual Report 2017–2018



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Annual Report Website: www.aavso.org/annual-report

On the cover...

Top: Attendees at the 2018 AAVSO Annual Meeting, Lowell Observatory, Flagstaff, AZ

Bottom: (clockwise from upper left) Stella Kafka; Bob Stephens, Kristine Larsen, Stella Kafka; Bert Pablo, Matt Craig, Richard Berry; Percival Lowell



Picture credits

In addition to images from the AAVSO and its archives, the editors gratefully acknowledge the following for their image contributions: Bob Stephens, Barbara Harris.

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1. The AAVSO in a nutshell

The AAVSO is a 501(c)3 non-profit association, whose mission is to enable anyone, anywhere to participate in scientific discovery through variable star astronomy. As such, the AAVSO is building an international community of Citizen Astronomers (professional astronomers, educators, students, non-professional observers) who are working collectively to acquire data, study, and understand aspects of variable stars. The AAVSO is a membership organization, generously supported through memberships and donations by its international community.



Member-observer Barbara Harris

The main aspects of the AAVSO's program are:

- The AAVSO International Database: a repository of variable star data, populated by
 observations from the AAVSO international community. Data acquired by visual, DSLR,
 CCD, PEP and solar observers of more than 25,000 variable objects are currently available
 to the research community through this database. Relevant data quality control is
 conducted by AAVSO staff and volunteers.
- The AAVSO's International Variable Star Index (VSX): this is a super-catalogue of known variable star information for more than 500,000 known variable stars. Objects in the catalogue include survey variables and new stars discovered by AAVSO observers.
- AAVSO alerts, campaigns, and fostering a close collaboration with the international
 astronomical community: the AAVSO builds connections between professional and nonprofessional astronomers, ensuring that AAVSO observers participate meaningfully in
 international collaborations resulting in cutting-edge science.
- AAVSO observer support: AAVSO software provides tools (such as finding charts, comparison stars, visible targets, etc.) for observers to plan and execute observations of appropriate targets. A helpdesk at the AAVSO headquarters provides answers and feedback to observers as needed. AAVSO-provided software also encourages the AAVSO community to engage in data processing and analysis.
- AAVSO educational/training activities: these aim at introducing various observing techniques, software, and the physics of variable stars.
- An active peer mentoring program: this program connects experienced observers with novices to provide guidance on aspects of variable star observations.
- The Journal of the AAVSO: a refereed publication that welcomes scientific, technical, and educational manuscripts by both professional and non-professional astronomers alike.

Through our work, our observing sections, our meetings, and our forums the AAVSO is building an international collaboration of individuals who are interested in understanding some of the most dynamic and serendipitous phenomena in the universe.

AAVSO Officers, Council Members and Headquarters Staff

You may contact these persons through AAVSO Headquarters.

Officers

Director	Stella Kafka	(term of office 1 February 2015—)
President	Kristine M. Larsen	(2015–2018)
1st Vice President	Bill Stein	(2017–2018)
2nd Vice President	Kevin B. Marvel	(2015–2018)
Secretary	Gary Walker	(2009–2018)
Treasurer	Robert Stephens	(2017–2018)

Council Members

Richard Berry	(2016–2018)
Tom Calderwood	(2016–2018)
Michael Cook	(2017–2019)
Joyce A. Guzik	(2015–2019)
Michael Joner	(2016–2018)
Katrien Kolenberg	(2014–2018)
Arlo Landolt	(2017–2019)
Gordon Myers	(2017–2019)
Gregory R. Sivakoff	(2016–2018)

AAVSO Headquarters Staff

Sara J. Beck	Technical	Assistant	(Science	Operations),
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Special Projects

Stella Kafka, Ph.D. Director

Sebastián Otero External Consultant, VSX Team, Spanish Translations Michael Saladyga, Ph.D. Technical Assistant, *JAAVSO*, *Newsletter*, and *Annual*

Report Production Editor

Kathy Spirer Operations Manager
Owen Tooke Administrative Assistant

Elizabeth O. Waagen Senior Technical Assistant (Science Operations),

JAAVSO Associate Editor, AAVSO Newsletter and

Annual Report Editor

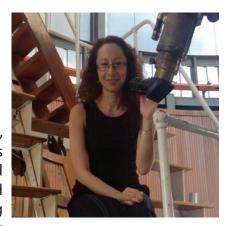


Annual Report of the Director for Fiscal Year 2017–2018

Stella Kafka, PhD, Director

A year in review

What a year! TESS was launched, Gaia DR2 was released, the Parker Solar Probe promises to reveal unknown aspects of the solar corona, the Kepler telescope ran out of fuel and left a legacy of thousands of unknown variable objects ... and we, at the AAVSO, have been busy focusing on completing projects that enhance our observers' resources, participating



in key collaborations bringing exciting projects to our community, and ensuring that our databases and software are operational and secure. As exoplanet discovery and exploration is becoming routine in the science world, we make sure that AAVSOers have access to educational material which helps them enhance their observing skills and acquire data for the targets that need it. Our CHOICE courses are improving thanks to the constant attention and hard work of our volunteer instructors, the JAAVSO's content is expanding to include new subject matters, VSX reached (and surpassed) half a million entries, we connect our observers with exciting projects through our alerts, and we focus on updating our software and database infrastructure with an eye to long-term sustainability and security. Thanks to our volunteers, we provide a valuable mentor program, translations of our manuals, updated information in our observing sections, charts, sequences, and data validation.

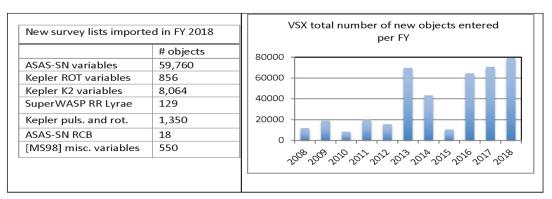
In addition to improving our science program, we are responsible for building maintenance, curating our archive and our library, ensuring smooth communication between HQ and our community, providing assistance to our observers when needed, ensuring that memberships are renewed, and updating our community about our activities. We have a strong presence in key conferences where both professional and non-professional astronomers discuss new science and upcoming projects, and we bring our community together through our meetings and forums. We always talk to our members and observers worldwide, and welcome constructive comments and contributions; after all, our mission is to serve this international community and we

need to understand its needs and ensure that we provide the necessary resources. All the above wouldn't have been realized without our community's support. Either through donations, memberships, or participation in NASA-funded scientific projects, funding has been secured which enables us to identify and hire the appropriate contractors to help us with Drupal migration, cyber security, IT, and software development. Our fundraising efforts enable us to re-design our servers, ensure frequent backups of our software and database assets, and maintain your favorite software. Because of your generosity, AAVSO staff work on data quality control, collaborate with volunteers, help new observers with their data, ensure our databases' quality, and curate more than half a million variable objects in VSX. The AAVSO is fortunate to have the support and encouragement of such an engaged and generous community, who enrich our databases with valuable data and help open the doors of variable star astronomy to all.

Here are some of the highlights from FY 2018:

Variable Star Index – improvements and updates

VSX is improving each year, enriching the database with new variable star submissions from individual observers and survey lists. In FY2018, we imported more than 80,000 new objects, totaling more than half a million objects in the VSX database. This is the most comprehensive (albeit not complete) database of updated variable star information, a unique and valuable resource to the research community. More on VSX can be found in the relevant section of this report.

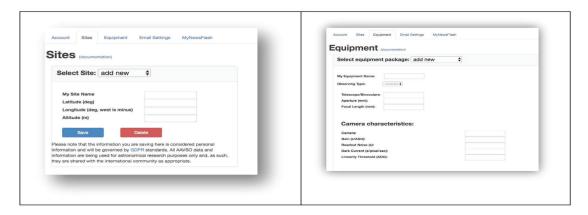


At the same time, we improve VSX's functionality based on our community's feedback. We generated a Frequently Asked Questions web page and we continue checking all new VSX submissions for accuracy, providing feedback when needed. Based on constructive forum comments, we assembled a team of volunteers from our international community to generate a comprehensive user's manual for VSX submissions, including examples and a detailed description of the material needed and relevant resources to

retrieve the material. The manual is considered to be a live guide – improvements will be implemented based on our community's feedback. We also established two new modes of for VSX submissions (beginner and advanced), including updates of existing records. In all cases, contributors will be acknowledged for their work on the VSX pages. The VSX manual can be found at https://www.aavso.org/vsx/

Exoplanet database (ExoDB)

A major new project launched in 2018 was our new exoplanet database, a project that was in the making for a couple of years. This project was triggered by discussions with professional astronomers who were interested in working with our community for exoplanet transit studies, and asserted that measurements of brightness variations vs time alone were not sufficient for their work. Thanks to the persistence and hard work of the exoplanet observing section leader, Dr. Dennis Conti, we designed a database that will ensure critical information on exoplanet transits is captured for long-term studies of their light curves and behaviors. The new database can be accessed through WebObs and was officially announced at the 2018 AAVSO annual meeting in Flagstaff, through a press release that was shared with hundreds of journalists around the world. More information can be found at: https://www.aavso.org/exoplanet-data



Observer Site & Equipment registration page

This is an online repository of information on observer equipment and observing site(s), and it provides an inventory of our observers' capabilities which can help the AAVSO provide appropriate targets for observers. Observing site information is also required for exoplanet data submission (as it is critical for timing information), therefore this new catalogue is required for the exoplanet database. This project was led and executed by volunteers, in particular George Silvis, working closely with HQ staff. Relevant

documentation accompanies the database, providing guidelines on how to fill out various fields in the forms.

Photometric Data Quality control updates

Ensuring that our light curves are of the high quality required for scientific projects is one of the priorities at HQ. Data quality assessment has been and is conducted by HQ staff, yet with the increase of data submissions, there was a need to recruit a team of volunteers who would help with the task at hand. To facilitate the team's work, "Zapper Validator", a software that helps identify discrepant data, was updated and disseminated to the team. We welcome more volunteer engagement in this team effort.

Updated "Standard Fields"

We added standard fields published by Arlo Landolt to the "Standard Fields" web page (https://www.aavso.org/apps/vsd/stdfields). Those additional fields enable our observers to determine transformation coefficients for their equipment at any time of year, as those multiple standard fields are spread out around the sky.

New Developer's Guide

To facilitate our collaboration with our volunteer and HQ programmers, and ensure we follow best practices that would lead to long-term maintenance and sustainability of our software, we composed a developer's guide. This handbook outlines guidelines for AAVSO software developers, outlining programming language requirements, software compatibility with our systems, and proper code documentation when producing software for the AAVSO.

Training material translations

A new translation of the DSLR manual in French is available; French and German translations of the Solar Observing Guide have also been completed and published. Translations have been made of the PEP guide in French, the exoplanet manual in French, and the web page "Variables: What are they and why observe them" in Uzbek.

CHOICE Courses

Thanks to the time and energy dedicated by our volunteer instructors, we continue offering our successful CHOICE course series, based on instructor availability. At the end of each course, we solicit participant course evaluations which are shared anonymously

with instructors – these help us identify challenges with the course material and improve course curriculum and format. Overall, the courses offered in FY2018 are:

Course Title	Instructor
CCD Photometry (Part 1)	Ed Wiley
Exoplanet Observing	Dennis Conti
Variable Star Classification & Light Curves	Blake Crosby
CCD Photometry (Part 2)	Phil Sullivan
Developing a Visual Observing Program	Michael Cook
How to use VStar	JoDee Baker-Maloney
DSLR Photometry	Mark Blackford

Technical Operations

2018 was a year of transformation and infrastructure upgrades. After performing an overview of our software, applications, hardware, back-end online security, and backups, we formed a plan to simplify our digital footprint, focusing on software that enables our community to plan, submit, and retrieve observations, projects with significant scientific yield, and a sustainable software structure that would enable better management of our databases and more pro-active upgrade of our tools. More details on our activities can be found in the "Astronomer's Corner". In short, we re-designed our live server and planned an identical development server which will help us test and upgrade software; we started migrating our web page from Drupal 7 to Drupal 8 and performing a necessary web page re-design and implementation (based on a user study); and we focused on improving and simplifying our backups, both on the AWC and on-site. Recognizing that cybersecurity is an issue for all organizations with an online presence, we started a system vulnerability assessment which will help us identify possible weaknesses of our firewall, which we will correct accordingly. Our best security assurance to our hundreds of observers is that we do not retain any sensitive financial information, and we are using our databases exclusively in a way that is appropriate to the mission of the AAVSO - sharing observations of scientific data, building collaborations between professional and non-professional members of our community, giving credit to our observers' hard work in scientific publications.

Outreach (FY2018)

- October 14, 2017: Astroassembly Skyscrapers (Rhode Island); talk: "The AAVSO Program: How to contribute to cutting-edge science"
- October 16, 2017: RASC Edmonton (Edmonton, Canada); talk: "Observing Variable Stars with the AAVSO"
- October 17, 2017: University of Alberta Physics (Edmonton, Canada); talk: "The AAVSO Program: A Resource for Variable Object Research"
- October 21, 2017: Remote presentation at the Hellenic symposium of amateur astronomy; title: "The AAVSO Program: How to contribute to cutting-edge science"
- October 24, 2017: Online panel participation for high schools; topic: "Extra-terrestrials or a Zombie Apocalypse What are the odds?"
- Week of November 13, 2017: IAU S339: Southern Horizons in Time-domain Astronomy (Stellenbosch, SA). Organized workshop on "Stellar variability: From citizen science to citizen astronomy"; Conference public talk: "Citizen Astronomy in the era of large surveys"
- November 22, 2017: Astronomical Society of Southern Africa (Cape Town, SA); talk: "Observing Variable Stars with the AAVSO"
- January 23, 2018: Concordia University Distinguished speaker; talk: "Studying Variable stars with the AAVSO"
- January 26, 2018: San Diego State University; colloquium: "The AAVSO Program: A Resource for Variable Object Research"
- March 8-11, 2018: Astronomical Society of Southern Africa Symposium 2018 (Cape Town, SA); talks: "The AAVSO for your research in the large surveys era" and "Citizen Astronomy in the Era of Large Surveys"
- April 5-8, 2018: Southern Star Astronomy Convention; talks: "Variable Stars and their Stories" and "The AAVSO Program: A Resource for Variable Object Research"
- April 10, 2018: Alachua Astronomical Society (Florida); talk: "Variable Stars and their Stories"
- April 11-13, 2018: University of Salisbury (Maryland); talk: "The AAVSO Program: A Resource for Variable Object Research"
- May 10-12, 2018: Pisgah Astronomical Research Institute (PARI; Ashville, NC); talks: "The AAVSO as a resource for Student Variable Star Research" and "Variable stars and their stories" (public talk). Also, multiple online presentations for summer interns at PARI in June and July.
- June 25, 2018: "Trimble-fest" (Irvine, California); symposium to honor Virginia Trimble's 75th birthday and 50th anniversary of her graduation from Caltech. Presented the AAVSO William Tyler Olcott Award for Distinguished Service to Virginia.
- July 22-28, 2018: RTSRE 2018 (Hilo, Hawaii); Keynote presentation: "The AAVSO as a resource for student research"

Treasurer's Report October 1, 2017–September 30, 2018

Robert Stephens, Treasurer, AAVSO, 49 Bay State Road, Cambridge, MA 02138

Audited Financial Statements

American Association of Variable Star Observers Statement of Financial Position September 30, 2018

Assets	
Current Assets	\$ 798,883
Cash and cash equivalents Prepaid expenses	\$ 798,883 5,223
Investments	12,903,911
Total Current Assets	13,708,017
Property and Equipment, Net of Accumulated Depreciation	1,265,009
Total Assets	\$14,973,026
Liabilities and Net Assets	
Current Liabilities	\$ 45,759
Accounts payable and accrued expenses Prepaid membership dues and meetings	\$ 45,759 53,299
Total Current Liabilities	99,058
Long Term Liabilities	
Security deposit and last month's rent re Annex	8,332
Net Assets	
Unrestricted	11,296,460
Temporarily restricted	201,382
Permanently restricted	3,367,794
Total Net Assets	14,865,636
Total Liabilities and Net Assets	\$14,973,026

continued on next page

American Association of Variable Star Observers Statement of Activities and Changes in Net Assets For the Year Ending September 30, 2018

Changes in Unrestricted Net Assets Unrestricted Revenues, Gains, and Other Support	
Contributions and grants	\$ 199,382
Investment interest and dividends	1,223,671
Membership dues and fees	76,649
Sales of publications and related material	9,941
Meeting fees Annex rent	22,149 50,000
Unrealized losses on available-for-sale securities	(911,973)
Gain on sale of investments	338,680
dam on sale of investments	
Net assets released by restrictions	1,008,499 26,103
·	
Total Unrestricted Revenues, Gains, and Other Support	1,034,602
Expenses	
Program Services - research, data collection, publication and meeting	s 630,347
General and administrative	403,539
Fund-raising	23,258
Total Expenses	1,057,144
Increase in Unrestricted Net Assets	(22,542)
Changes in Temporarily Restricted Net Assets	
Contributions and grants	8,275
Investment interest and dividends	14,542
Unrealized gains (losses) on available-for-sale securities	(11,916)
Gain on sale of investments	4,425
Assets released from program restrictions	(26,103)
(Decrease) in Temporarily Restricted Net Assets	(10,777)
Changes in Permanently Restricted Net Assets	
Contributions	1,750
Increase in Permanently Restricted Net Assets	1,750
Increase in Net Assets	(31,569)
Net Assets—Beginning of Year	14,897,205
Net Assets—End of Year	\$14,865,636

Variable Star Observing Campaign Highlights

Elizabeth O. Waagen

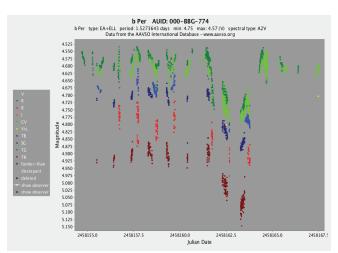
The AAVSO participates in many observing campaigns on variable stars. These campaigns arise from the request for assistance by an astronomer, or from the AAVSO itself in response to a discovery or other unusual stellar activity. Campaigns may run from a few days to weeks or months, or may even be ongoing for several years.

Observing campaigns carried out in FY 2018 by the AAVSO on established or suspected variable stars and on the 14 galactic novae discovered during the year may be read about via the AAVSO webpage on AAVSO Alert Notices for Observing Campaigns and Discoveries (https://www.aavso.org/aavso-alert-notices-for-observing-campaigns-and-discoveries). Note that numerous campaigns from earlier years listed on this webpage are still underway.

Below are a few highlights from the observing campaigns in which the AAVSO was involved this year.

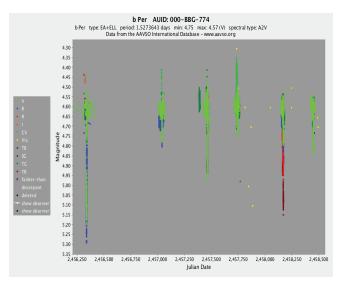
b Per (HD 26961)

The bright star b Per (HD 26961) continues to be of great interest. Two campaigns were carried out this year on this triple system - one beginning in January (AAVSO Alert Notice 610) and one beginning in October (AAVSO Alert Notice 655) - at the request of Dr. Donald F. Collins (AAVSO member, Swannanoa, NC) and colleagues Dr. Robert Zavala (US Naval Observatory, Flagstaff Station), Dr. Anatoly Miroshnichenko (University of North Carolina at Greensboro), and Jason Sanborn (Lowell Observatory).



The campaign in AN 610 was on the third star in this triple system, which was predicted to transit in front of the inner pair of stars (a primary eclipse) during the week centered on February 12, 2018, and time series photometry was requested. In AN 655, the target was also the third star, which was predicted to transit behind the inner pair of stars (a secondary eclipse) during the week centered on November 21, 2018, and time series photometry was again requested. Both campaigns

were successfully concluded as AAVSO observers provided multicolor photometry that



showed the eclipses on February 12-14 UT (Figure 1) and November 20-22 UT, respectively.

These campaigns were the fifth and sixth in a series with AAVSO participation to observe transits among the three stellar components of this system in order to assist in untangling the complicated orbital relationships. AAVSO observers have observed six transits in this system (Figure 2), including the first recorded eclipse, which occurred in February 2013. Links to more information about this system and an animation, as well as

to the four earlier AAVSO campaigns, may be found in the Alert Notices.

ASASSN-18ey (= MAXI 1820+070)

A new black hole X-ray binary candidate, ASASSN-18ey (= MAXI 1820+070), was discovered on March 11, triggering an immediate flurry of ground- and space-based professional observational campaigns. On March 19, Drs. Gregory Sivakoff (University of Alberta), Poshak Gandhi (University of Southampton), and Diego Altamirano (University of Southampton) requested AAVSO assistance in providing optical photometry in support of some of these campaigns and of additional ones being planned (AAVSO Alert Notice 624). Two days later, Dr. Tom Maccarone (Texas Tech University) asked for AAVSO observations in support of his VLT (ground-based, visible and infrared) and XMM (space-based, X-ray) observations (AAVSO Alert Notice 625). A month later ASASSN-18ey had stalled in its X-ray state evolution, and a multiwavelength "blitz" campaign occurred, with several facilities observing the binary over a 24-hour period; the AAVSO was asked to participate (AAVSO Alert Notice 630).

ASASSN-18ey continued to evolve over the following months, and AAVSO observers continued to monitor it faithfully. By late December, when the observing season ended, ASASSN-18ey had finally faded to fainter than V magnitude 17.8 and AAVSO observers had provided over 372,000 multicolor observations. Several astronomers are using AAVSO data in their analyses of this fascinating system.

FO Agr

Beginning in July 2018, a campaign (AAVSO Alert Notice 644) was carried out on FO Aqr, an intermediate polar cataclysmic variable featuring a magnetized white dwarf that accretes from a low-mass companion star. This type of star has "high" (bright) and "low" (faint) states; the method of mass transfer is believed to change depending on the state.

When FO Aqr entered its low state in July, Drs. Colin Littlefield (University of Notre Dame), Peter Garnavich (University of Notre Dame), and Mark Kennedy (University of Manchester) requested AAVSO optical photometry until FO Aqr returned to its high state. Dr. Littlefield wrote: "AAVSO observations enable us to identify which periodicities are present in the light curve, and from that information, it is possible to draw conclusions about FO Aqr's accretion mechanisms. In particular, we are interested in using AAVSO data to test the hypothesis by Hameury & Lasota (2017) that FO Aqr's disk dissipates if a low state is deep enough."

This campaign was the third one on FO Aqr for these astronomers, with previous ones occurring when FO Aqr entered its low state in 2016 (first low state ever observed for this star; AAVSO Alert Notice 545) and in 2017 (AAVSO Alert Notice 598). AAVSO observers followed FQ Aqr closely through the observing season, and detected clearly when it returned to its high state in December. Two papers have been written to date that include analysis of the 2016 and 2017 AAVSO data, with AAVSO observers as co-authors.

Astronomer's Corner

Bert Pablo, Staff Astronomer

Change is on the horizon at the AAVSO. Some of it is tangible, like the new face our website will soon have, but much of it is below the surface. So as we gear up for what will definitely be an exciting 2019, let's take a look at everything that has been done over the past year.

First let's take a look at the new capabilities we have and are still adding for our users. We have officially released two new pieces of software. The first is the Site and Equipment Page. This allows users to put in information about the location and make up of their instrument packages. While this information is useful to keep track of for all of your observations, it is a necessity for our new programs and databases. As you have undoubtedly heard the AAVSO is officially in the exoplanet game with the launch of our new web application ExoDB. This tool provides a new way of submitting time series data, specifically geared towards exoplanet science.

Right on the heels of ExoDB is the long-awaited debut of our spectroscopy database. While it is not quite officially released, the application is already in beta testing and is almost ready. Since spectroscopy is very different in many respects from the photometry our community is accustomed to, we have been working diligently to construct new observing materials and manuals. This will provide our whole community the tools necessary for exploring variable stars in a new and different way.

Beyond our new applications much of this year has been spent redefining our technological footprint. Now you might be wondering why need to put effort into this, and it's a fair question. Unfortunately, despite the effort and time involved in such a project it's a necessary evil. The internet and how we access it is constantly changing, which means all of the programs we use and create must necessarily change as well. This invariably will cause issues, issues we could not foresee, or take steps to mitigate, within our current architecture. Therefore, we either kept things the same, working around bugs that would crop up, or had to interrupt service to our community to fix them. In order to make this process more seamless we are moving away from a "one server for everything" approach and segmenting our infrastructure into smaller more manageable chunks that will be both more efficient and more sustainable moving forward. This new design is being implemented as we speak, and will be the platform for our updated website.

Speaking of updates, the rebuilding and re-inventing of our infrastructure has allowed for a much needed update to all of our servers internal systems, from the operating system to Drupal, the website container itself. While much of this you won't notice, the changes to the website will be hard to miss. Due to fundamental changes in how Drupal behaves and interacts with other programs this necessitated some large rewrites of the underlying code, which will change the look and feel a bit from what you're used to. It will still be the same website content, just displayed in a slightly different way, which we hope will make your time on our website easier.

Finally, I wanted to make you aware of all the effort we are putting into security. One of the main reasons for updating our infrastructure is so that we are able to patch security vulnerabilities that are found with time. In the past we have been reactive, dealing with problems as they arose. However, that isn't a position anyone wants to be in and we are putting substantial effort into becoming more proactive. In addition to updates and infrastructure changes we are currently undergoing a full technical security assessment, so that we can be made aware of the areas that are problematic and can address them in due course. We take pride in being an organization you can trust with your legacy and we are doing everything in our power to stay that way.

The International Variable Star Index (VSX)

Sebastián Otero and Patrick Wils

This report covers activity from October 1, 2017 to September 30, 2018.

The International Variable Star Index is a continuously evolving database including most available variable star catalogues and stars published in variable star listings/journals plus new discoveries submitted by individual researchers.

It was created by amateur astronomer Christopher Watson and is kept by the AAVSO with the valuable input of individual contributors who revise known variable stars and submit new ones. All that information is included in our database after a strict moderation process.

VSX is mostly maintained by staff member Sebastián Otero and volunteer Patrick Wils. Sebastián reviews all submissions/revisions and contacts the observers so they can correct or improve the uploaded information. Questions about catalogues and data analysis -and especially issues concerning variable star classification - are continuously being discussed by e-mail as part of the moderation process. He also adds stars from alert pages, works on updates of existing records, prepares new lists of objects for upload, hides duplicate entries and works on inconsistencies between VSX, VSD (The Comparison Star Database) and the AID (AAVSO International Database). He also deletes variable comparison stars from sequences (the Sequence Team then replaces the "lost" comparison if needed).

Patrick works with data updates from journals and some other miscellaneous sources. He uploads lists of variable stars prepared by Sebastián, Klaus Bernhard and himself, helps replacing AUIDs in VSX when an already existing AUID in VSD is found and does some of Sebastián's duties when he's not available. He is always behind the scenes making changes and improvements to the database structure and the user interface and correcting bugs, things that may go unnoticed but make the VSX process faster and more efficient.

Below are some of the changes he implemented:

At the beginning of the fiscal year we added a VSX Frequently Asked Questions (FAQ) page. Information that had to be propagated through individual e-mails is now being presented there. Information on data-mining, reference formats, photometry and passbands (including online spreadsheets to transform and correct data), identifiers, supporting documentation and submissions in general can be found in this new

resource.

- Catalog data from 2MASS, AllWISE, GALEX and APASS have been added when available in a new section of each star's detail sheet.
- Sky coverage plots can now be displayed.
- External links to ASAS-SN and Pan-STARRS1 data are now available

Volunteers

Volunteers play a major role in keeping VSX up to date and working smoothly. Patrick Wils has dedicated a lot of volunteer work over these years to add new variables or corrections made on known variables as they are published in journals, alert pages or even in web resources.

Klaus Bernhard continued helping us preparing lists for VSX import. It is not something straightforward because each paper or each variable star list comes with its own format and we need to extract the information in a format suitable for our database needs. A very important step in this work is to make sure that the new variables added are not actually duplicates of stars already included in VSX. So a thorough cross-identification of the new stars with our own records is always performed. We check the reliability of the published data so we can avoid overwriting good information with wrong numbers in the case of revisions and we make sure that we are adding bona-fide new variables in the case of the new entries. Patrick Schmeer has been helping us adding new objects from the alert pages or updating the information on existing ones.

We need more volunteers. If you have experience with variable star classifications, light curves, and if you have used VSX in the past, you might want to give these tasks a try:

- Checking Astro-ph papers;
- Checking various online journal papers;
- Preparing lists for VSX upload;
- Revising VSX records with obvious mistakes or missing information.

If you would like to join us, send a message to vsx@aavso.org

New Variable Types Added

We always try to update our variability types document by adding the most recent variable star types recognized in the literature.

The EL CVn (EL, systems consisting of a core helium-composition pre-white dwarf and an early-type main-sequence companion) and HW Vir (HW, systems consisting of a hot

subdwarf and a red or brown dwarf companion displaying reflection effect) subtypes of eclipsing binaries were added.

We also added the SN Ic-BL (Broad-lined SN Ic supernovae, sometimes called "hypernovae") subtype of supernova.

Number of Submissions and Revisions

- We've had 1084 new variables submitted to VSX by individual users (other than VSX Team members) this year (860 in 2016-2017).
- The number of revisions made by users more than doubled from 239 last year to 495 in 2017-2018.
- We currently have 376 different users that have submitted at least one submission or revision to VSX. 28 of them had their first VSX experience this year.
- The VSX Team updated data on 29468 known variable stars (3377 last year).
- The number of new objects added by our team keeps growing each year. For the fiscal year 2017-2018 we added 78272 new objects, while last year they were 69877.
- This means a total of 79356 objects (adding the 1084 objects submitted by individuals).

You can check what's new on VSX by trying one of the special searches (like "Changes since last login") in the VSX search page.

VSX Submission Policy

Even though it hadn't been implemented yet by the end of the fiscal year 2017-2018, it is important to announce that a new Novice discovery checkbox has been implemented in the New Star Form and New Star Wizard so VSX users that do not have the interest to use survey data or make a deep analysis of their data can still submit their findings to VSX. We encourage full analysis but if someone has a random finding and doesn't feel like getting too involved in variable star analysis (e.g. asteroid observers), he/she can now submit their discoveries to VSX.

VSX Forum

We suggest all VSX submitters to use this forum for pre-submission discussions so everybody can share their experiences and learn from one another. There are very experienced observers among the VSX users so having them sharing a common space allow people to discuss their submissions, ask questions, post doubts and eventually get to the submission stage with a higher quality product.

Please keep using the forum in 2019. It is at: https://www.aavso.org/forums/about-aavso/vsx. And remember that to start discussions in the VSX Forum you have to create an AAVSO website account: https://www.aavso.org/apps/register/

Duplicate Records

VSX has currently over 600,000 records. We don't call them stars because there are still some duplicate records among them but we have been working hard with to minimize that problem over these years.

Hiding visible duplicates helps avoiding confusion when an observer finds two stars at nearly the same position and can't decide if there are actually two variable stars there or they are just one and the same. Software can be fooled by these duplicates too and our International Database may suffer the consequences with spurious reports being submitted. We surely don't want that!

In the framework of this primary record creation work (which means that all the information available is used to update a star's detail sheet), Sebastián hid 179 duplicate visible entries this year plus 88 unclassified (not visible for the public) duplicate objects. 6,553 duplicate records were hidden since the primary record creation work started back in 2011 (6,835 counting the unclassified ones). Patrick hid another 112 records this year after cross-identifications were made while importing new lists. This will always be work in progress. A total of 28,443 objects have been hidden since VSX was launched in 2005.

Incorrect Identifications Corrected

More incorrect identifications are being found in the process of cleaning up the VSX database.

- 34 incorrect cross-identifications in VSX have been corrected this year (usually incorrect identifications made by surveys).
- 58 GCVS/NSV identifications have also been corrected.

It is nice that observers report incorrect identifications when they find them. Let us know if you find any.

Cross-Identifications (between existing entries) added

208 new cross-identifications between VSX records were established this year (2988 in total since 2011) and the resulting duplicates were deleted.

Work on VSX/AID: inconsistencies and problems with submitted data

Work to clean up the AAVSO International Database (AID) from errors caused by duplicate entries in VSX (most of them not visible to the public) has also continued. The fact that we have the same object listed two or even three times under different names and with different AUID's means that we may have data for the same object split in different records. Thus, the light curves are not as complete as they could be.

We individually corrected 124 such records in 2017-2018. Sara Beck moved all the observations to the new primary records and I deleted the AUID of the obsolete records so people can't submit data using the old identifiers anymore.

As usual, we have also contacted several observers to modify wrong observations reported to the AID that are found while analyzing AAVSO data to improve the information delivered in VSX. We urge observers to double-check their images to properly identify the stars being reported.

Remember that we periodically update the list of stars with companions that may cause identification or photometry problems. It can be found here: https://www.aavso.org/variable-stars-companions

Work on VSX/VSD: comparison stars that turned out to be variable

Finally, we don't want variable stars to be used as comparison stars, but this may happen sometimes.

There were not enough data some years ago to judge if some stars were variable or not and they could have been selected as good comp stars based on color or proximity. Now, with more survey data available or with observations provided by our observers, we can identify that some of those comp stars are actually variable. Work is being done to eliminate these stars from our sequences and find suitable replacements. 10 stars have been eliminated this year.

VSX is a core application that interacts with almost everything else in the AAVSO universe, from other software tools to the observers submitting data via WebObs. We try to improve it every day, solving inconsistencies and updating the database with the most recent data available.

We thank all the people who contribute to VSX and all the AAVSO staff that helps in the cleaning up process.

The AAVSO Network of Remote, Robotically Controlled, and Automatically Queued Telescopes (AAVSOnet)

Arne A. Henden and Mike Nicholas

The AAVSO Robotic Telescope Network, AAVSOnet, started in 2005 with a single telescope, SRO35, located in Sonoita, Arizona. In 2008 we added OC61 by partnering with the University of Canterbury in New Zealand at Mt. John Observatory. 2009 saw the addition of the first Bright Star Monitor at Astrokolkhoz Observatory near Cloudcroft, New Mexico. In 2018 there were eight active telescopes: the BSM systems (NM, South, Hamren, Berry, and NH), SRO50, TMO61, and OC61.

Each of these telescopes uses identical software: MaximDL for image acquisition; ACP and ACP Scheduler for controlling the telescope, camera and scheduling observations, and FocusMax for focusing each system. The AAVSOnet Software Automation system (ASA), designed and implemented by George Silvis, is used to maintain and manage jobs sent to the telescopes using comma-delimited EXCEL files. Each telescope is either at a University or located at a private site. Volunteers perform any local maintenance and operation. AAVSO members, as a free benefit, can propose to observe specific targets. These proposals are reviewed by a small Telescope Allocation Committee. Those approved are then put on the telescope queues. When images are taken, they are transferred back to HQ where they are automatically dark-subtracted and flat-fielded. Processed images are then placed on the AAVSO ftp site, and/or uploaded to the VPhot cloud analysis program at the discretion of the individual investigator. In the background we also extract all stars in every image to be placed into the on-line Epoch Photometry Database at a later date.

This year BSM_NH has under gone a name change to BSM_NH2 with the reconfiguration of its system to include a higher accuracy Paramount ME mount, loaned by Tzec Maun, a new ZWO ASI183mm-pro-cool CMOS camera and filter wheel and a new rollup enclosure, designed and built by Gary Walker. BSM_NH2 has been operational since September and is producing higher-quality data than the original system, with better tracked images and finer angular resolution. The original ST8-XME camera will be sent to BSM_Berry to replace the one loaned by Greg Bolt.

BSM_Hamren is moving to a drier site on the main island of Hawaii and will have an unobstructed view of the horizon. Mike Linnolt will continue to host the BSM_Hamren and is working with Gary Walker on setting up the new site with a new automated enclosure designed by Gary. A replacement for its failed camera has been donated by Dick Post. These changes will significantly improve the operation and reliability of this

valuable site. It is anticipated that BSM_Hamren will come on line in early 2019.

A new BSM system, BSM_TX, is being configured and should be ready for operation shortly. The site is located in west Texas. This system is owned by Dave Cowall, who is providing it for use by the AAVSOnet community. It is expected to be on-line in the spring of 2019. This will bring the total number of AAVSOnet sites to nine.

The original headquarters computer server (OCCAM) was replaced with a more secure and updated system. The new system has taken over the AAVSOnet image pipeline. A second identical server was also added as backup. The cost of both systems was paid for by Dick Post and was installed with the help of George Silvis. The AAVSOnet image pipeline software has been improved and made more reliable, thanks to the continued efforts by George Silvis, with help from Cliff Kotnik.

We continue to enjoy the support of many AAVSO member volunteers who manage, maintain and repair the AAVSOnet resources. Many thanks to Robert Dudley, Damien Lemay, Ken Menzies, Lou Cox, Jean-Bruno Desrosiers, Dave Hinzel, Duane Dedrickson and JoDee Baker for performing this valuable service. In addition, we'd like to thank our site managers and operators as well: Helmar Adler, Arne Henden, Nigel Frost, Bart Staels, Jon Holtzman, Dirk Terrell, Walt Cooney, John Gross, Peter Nelson, Greg Bolt, Mike Linnolt, Gary Walker and Bill Stein.

A web-page for the new BSM Section was created by Mike Nicholas, with help from Owen Tooke. It focuses on the BSM resources and provides members details on the equipment and how to propose their projects. Mike is the page Administrator, with help from Arne Henden, the Science advisor, and Ken Menzies, the Technical Advisor.

During 2018, a total of 15 new proposals were accepted, from professionals as well as amateurs. These included individual research, monitoring of objects for campaigns, and time series observations for some professional members of the AAVSO. The AAVSOnet telescopes collected 161,122 images of over 2,000 targets. Many of these are part of the BSM survey of all variables brighter than 8th magnitude.

The AAVSO Photometric All-Sky Survey (APASS) Arne A. Henden

APASS started in late 2009 in the north, and about a year later in the south. The goal is to cover the entire night sky, with every object being observed on at least four photometric nights. The main survey covers the magnitude range 10<V<17 in the Johnson B&V and the Sloan g'r'i' passbands. The expected final astrometry will be within 150milliarcseconds; the photometry should be better than 0.02mag for bright objects. This catalog was designed to do for photometry what the positional catalogs (such as UCAC and USNO-A) did for astrometry: provide calibrated references in every CCD field of view.

The original survey was funded by the Robert Martin Ayers Sciences Fund. In 2014, the NSF awarded the AAVSO a 2-year grant to both complete the observations as well as produce a final catalog. As part of the NSF proposal, we are extending the catalog with a Bright Star Extension, covering the range 7<V<12 and with BVu'g'r'i'zY passbands. In 2017, we asked for and received a one-year no-cost extension to the original NSF grant, and later received another extension to July, 2018.

The equipment at each site is composed of two ASA N8 20cm astrographs, Apogee Aspen CCD CG16m (KAF-16083 sensor) cameras and filter wheels, coaligned on a Paramount ME. The northern system is located at Weed, NM; the southern system is at CTIO in the MEarth roll-off building.

To date, over 530,000 images have been taken on about 1500 nights (combined north and south). Ten data releases have been made, with the most recent one occurring in October, 2018. In DR10, a total of 128 million objects have a minimum of two observations each, covering about 99% of the sky. DR10 was a complete reprocessing of all images, using Sextractor for star-finding and centroiding, astrometry.net and custom plate solutions, DAOPHOT aperture photometry, and a new collation technique to find measurements of the same star. This completed the processing of northern and southern images taken through 2016. The APASS "means catalog" can be searched online at the AAVSO web site. The Epoch Photometry (individual measures) can also be searched on-line if you are an AAVSO member.

Doug Welch (McMaster University) has been the key person in the data reprocessing, gaining access to the Canadian SHARCnet parallel processing computer, and providing storage for all images and star lists. Brian Kloppenborg has created the new collation program, which has eliminated many of the inconsistencies of the original algorithm.

APASS is being used by many groups worldwide, and is also being used by individual researchers for obtaining precise photometry of their favorite targets. We get over a dozen requests annually from professionals who want access to the entire catalog. Within the AAVSO, APASS is being used primarily for the generation of photometric sequences around program stars, and photometric confirmation of new submitted variable stars to VSX. The VisieR group in France hosts DR10 publicly.

We are continuing observations with the APASS sites to fill in a few missing fields, and processing about a hundred additional nights of earlier data that were missing from DR10. Additional data releases will be made as these data are analyzed.

A great many people have been involved in the APASS development. The PI of the project is Arne Henden; Dirk Terrell has provided computers, software and analysis; Stephen Levine is the primary astrometry expert; Doug Welch is archiving all images and photometry, serving catalogs and performing the Sextractor and astrometry.net initial processing; Brian Kloppenborg creates the merged catalog; and Ulisse Munari is providing quality control and external comparisons. In addition, there are a large number of volunteers, staff and students, including at least: Tom Smith, Aaron Sliski, Alan Sliski, Ken Launie, Shouvik Bhattacharya, Anisha Sharma, Patrick Wils, John Gross, Bill Stein, Sebastian Otero, Matt Templeton, Doc Kinne and Sara Beck. The UNC group (especially Kevin Ivarsen and Josh Haislip) provided on-site support when APASS South was at Prompt; Jonathan Irwin has been enormous help since the system moved to MEarth, along with the CTIO telescope operators. We've also had equipment and software contributions from Tom Bisque (Software Bisque), Bob Denny (DC3 Dreams), Doug George (Diffraction Ltd.), Apogee CCD, and Don Goldman (Astrodon). We thank them all – without their support and help, this project would never have happened!

The Journal of the American Association of Variable Star Observers John R. Percy, Editor

The Journal of the American Association of Variable Star Observers (JAAVSO) is the peer-reviewed research publication of the AAVSO, dedicated to variable star astronomy and a wide variety of related topics. It exists to disseminate the scholarly work of AAVSO members, observers and others to the AAVSO "family" and to the astronomical community at large; to demonstrate the scientific value of AAVSO data and thereby motivate AAVSO observers; to record the scientific content of AAVSO meetings; and to inform and inspire our members and others about variable star astronomy. It demonstrates, among other things, that small observatories, and skilled amateur astronomers, and student researchers can continue to make significant contributions to variable star astronomy.

In 2017-18, we published two issues, with well over 200 pages of content in total, in our (relatively) new large-page format. We continue to use an automated manuscripthandling system, and "blind" refereeing. The published papers covered the usual range of variable star topics, including papers on research with new data and old; instruments, methods, and techniques; data; education and outreach; history; book reviews; and abstracts of papers presented at AAVSO meetings. JAAVSO authors come from countries around the world, and include professional astronomers, amateur astronomers, and students. A typical issue contains about a dozen research papers and two dozen abstracts. This year, we began to publish short review/update articles, beginning with one on Young Stellar Objects by Professor Bill Herbst. Others are in the offing, and we are open to publishing either long or short reviews/updates on topics which are relevant to our readers. I have also continued to provide editorials on topics which I hope are both relevant and interesting. I would be happy, at any time, to receive suggestions about any aspect of JAAVSO, including possible topics for future editorials or review articles. One small concern: the range of topics of the research papers in recent issues has been rather narrow. We encourage papers on the observation, analysis, and interpretation on all types of variables, as well as on the topics mentioned above.

In late 2017, I surveyed the Editorial Board and Headquarters staff about the present and future of JAAVSO. Please see my editorial in Volume 45, Number 2. The AAVSO Council has struck a small committee to discuss and recommend on possible future directions for our flagship publication. It has also embarked on a search for a new Editor, since I do not plan on serving in that capacity forever.

As always, I am grateful to the Editorial Board for their advice and assistance, to the many voluntary and anonymous referees who ensure the integrity of the JAAVSO content, and to the Headquarters staff, especially Michael Saladyga and Elizabeth Waagen, for their hard work and excellent judgement in carefully editing and attractively formatting over 200 pages of informative but complex material.

JAAVSO Editorial Staff and Editorial Board

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Volumes of the JAAVSO can be found at https://www.aavso.org/apps/jaavso/volumes/

Awards 2018

AAVSO Merit Award Recipient

Barbara G. Harris

"...in recognition of her years of dedicated service to the organization, selflessly contributing in numerous ways, including as a member of Council and multiple Committees, mentoring fellow observers, teaching the CHOICE course in DSLR Photometry,



submitting over 38,000 variable star observations using multiple techniques, and in particular her valuable contributions to observing campaigns, including her discovery of the January 28, 2010, outburst of U Sco."

William Tyler Olcott Distinguished Service Award Recipient

Virginia Trimble

"...for her influential five-decade-long body of research in astronomy, including work on myriad types of variable stars; her outstanding service to the discipline through her annual reviews of advances in astrophysics as well as her work with professional organizations; her ability to bring to life the history of astronomy and communicate effectively with the public; and her invaluable service as a mentor and role model to multiple generations of astronomers."



AAVSO Director's Award Recipient

Dennis Conti

"...for his excellent leadership and work with the AAVSO Exoplanet program, training observers through the relevant CHOICE course, workshops, and manuals, creating and overseeing the design of the AAVSO exoplanet database, and his continuous involvement in cutting-edge science with HST and TESS.



Special Awards presented at the AAVSO 107th Annual Meeting

The AAVSO extends its grateful appreciation and sincere thanks to

Tyler Arnold, Zachary Burkhart, Keanu David, Prof. Carolyn Shiery

"for their valuable volunteer contributions to the AAVSO through their work on creating new marketing material promoting the AAVSO's work to the community."

The AAVSO extends its grateful appreciation and sincere thanks to

Juan-Luis González Carballo, David H. Hinzel, Gabriel Murawski, Sebastián A. Otero, Patrick Wils

"for their valuable volunteer contributions to the AAVSO and the variable star community through their work on generating the VSX new data submission manual."

Section Reports

Bright Star Network (BSM)

Section Leader: Michael Nicholas

The AAVSO Bright Star Monitor global telescope network has had a good year:

- 1. The BSM telescope network has implemented 6 proposals in the last three months of 2018.
- 2. A BSM web-page for the new BSM Section was created by Mike Nicholas and Owen Tooke, with help from the BSM Team. The page provides to members details on the equipment and how to propose their projects.
- 3. A poster, describing how the membership can use this free resource, was discussed at the Fall Membership meeting in Flagstaff.
- 4. The three BSM systems in the northern hemisphere have undergone modifications during the past 6 months:
 - BSM_NH2, New Hampshire, (Arne Henden/Helmar Adler) is coming back online at its new location, with a CMOS camera and a Paramount. The CMOS camera has already collected nightly images since October and appears to promise accurate photometry for brighter variables than previously possible.
 - BSM_NM, New Mexico, (Bill Stein/Dick Post) has seen the return and installation of its repaired/cleaned camera. The normal SW monsoon has ended, and the system is now conducting active imaging. It is receiving a pipeline improvement for local calibration processing of science images to complement the HQ pipeline.
 - BSM_Hamren, Hawaii, (Mike Linnolt) is currently inactive due to a camera failure. A new enclosure (constructed by Gary Walker) will be sent to Hawaii to replace the current "tarp" and the system will be moved to a drier location (owned by Mike L.). A replacement camera donated by Dick Post has been provided. This upgrade should provide a significant improvement in reliability and an increase in imaging. Hopefully the new system will be on-line in early 2019.
- 5. The two BSM systems in the southern hemisphere (Australia) continue to work well, under the operation of Greg Bolt and Peter Nelson. Greg Bolt has loaned his own camera for use at BSM_Berry, while a replacement camera for one which failed is being obtained.
- 6. The following table provides a summary of the number of nights where weather cooperated and science images were taken during the year:

System	Nights Run in 2018
BSM_NH2 ¹	110
BSM_NM	73
BSM_Ham	42
BSM_S ²	99
BSM_Berry ³	149

1Combined nights from BSM_NH and the newer BSM_NH2 site. 2 "BSM_S operates every clear night. At this site this equates to approximately 30-40% of the year. Sometimes it will run all night, others only several hours." - Peter Nelson. 3BSM_Berry takes the prize where in one, 9-hour night, 66 targets were observed and 430 science images were recorded.

7. Finally, we have received an additional BSM network site through the generosity of David Cowall. Dave is offering the use of his 101mm refractor system in West Texas at no cost to the AAVSO and will permit the BSM Section to conduct ACP - scheduled imaging projects. Dave will be the "operator" and Frank Parks (new member) will be the "image reviewer". George Silvis is in the process of making the web connection and Arne Henden, Ken Menzies, Gordon Myers and Helmar Adler will help with setting up the calibration and image pipeline. Bob Denny, DC3 Dreams, will donate three Sloan filters to the existing list.

BSM_TX will consist of the following equipment:

Hardware:

- Televue NP101 Telescope (#1291)
- Losmandy 12 Tripod Extension (UEQ-7989)
- Technical Innovations RoboFocus
- SBIG STF8300M camera (#83F180299)
- SBIG/DL FW8-8300 8 position filter wheel (B, V, R, I, clear filters)
- Losmandy G11 mount (#HGM9606048) with Gemini II upgrade
- Shuttle DS77U5 computer with SSD and DIMM cards installed
- Web Power Switch
- BK Precision Digital Display DC Power Supply

Software:

Windows 10, Maxim DL, ACP Expert, FocusMax

The membership will now have six BSM sites in 2019 for their observing and educational projects which should enhance the ability of the Bright Star Section to support our members and the mission of the AAVSO!

Charts and Sequences

Section Leader: Tom Bretl

The Charts and Sequences Team is made up of volunteers and staff who work behind the scenes to keep the Variable Star Plotter (VSP) as up-to-date, accurate, and useful as possible. The current members of the team are Arne Henden, Barbara Harris, Bob Stine, Bruce Sumner, John Toone, Robert Fidrich*, Keith Graham, Jim Jones*, Patrick Wils, Sara Beck, Sebastian Otero, Tim Crawford*, Tom Bretl*, Natalia Virnina, Brad Walter*, Matthew Templeton, Mati Morel, Michael Poxon*, Jack Davis*, Stella Kafka, and Elizabeth Waagen. (* members who submitted new or revised sequences)

The team continued to be very active this year. Here are some approximate numbers for January 1 through December 21, 2018:

- 555 new sequences
- 116 revised or updated sequences
- 65 CHET responses (some revisions, some new)
- 194 ASASSN sequences (some revisions, some new)

For comparison purposes, here are the totals from the last five years:

Year	New/revised Sequences
2018	687
2017	600
2016	597
2015	695
2014	259

The recent release of APASS DR10 has made it possible for the team to catch up on many past CHETs and sequence requests. More than 100 sequences were created, revised, or updated during just the first three weeks of December.

Thank you to all of the active team members who have pitched in to quickly fulfill requests for both new and revised sequences, and a very special thanks to Tim Crawford who has personally contributed almost 300 sequences during the past year. An average of approximately one sequence per day is a pretty amazing accomplishment!

Charts and Sequences links:

CHET submissions are made online at https://www.aavso.org/chet.
Requests for new sequences are made via email to compstars@aavso.org.
Complete instructions for doing so are given at https://www.aavso.org/request-comparison-stars-variable-star-charts.

Completed work is shared via a spreadsheet that is accessible at https://docs.google.com/spreadsheets/d/1mR4I7bEIFYZI5IwkkVEBwByCNXwiKCMzIPS1IAx0QvQ/edit?hl=e n&pref=2&pli=1#gid=317284472.

Every few months all observers are made aware of new and revised sequences via the AAVSO website News. Links to past updates can be found at https://www.aavso.org/category/tags/charts-and-sequences-updates.

The team website continues to be a work in progress at https://www.aavso.org/sequence-team-homepage, but it already provides access to a number of helpful files.

The sequence team is currently developing a new website where members can find things like SeqPlot instructions, sequence creation guidelines, and information about photometric resources other than SeqPlot. This should prove especially useful to new team members. At the moment the site consists only of links to relevant files, and is definitely a work in progress.

Eclipsing Binary

Section Leader: Gerard Samolyk

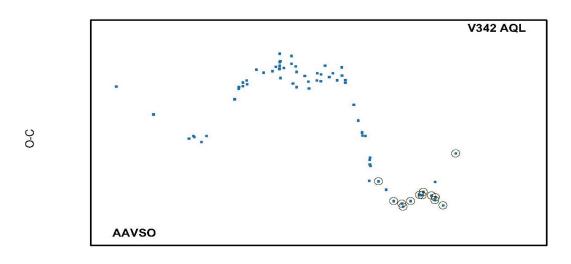
As of now, we have comparison stars for all of the stars on the AAVSO EB legacy program. Many of the other stars frequently observed by our group also have comp stars. I would like to thank the AAVSO chart and sequence team for their work in this area.

In 2018, two papers containing 774 times of minima of 313 stars observed by 18 observers were submitted to the JAAVSO in 2018. Observers who would like to contribute data to these papers in the future should upload their observations to the AID and send a copy to gsamolyk@wi.rr.com.

In the past year, 193 of the 201 stars on the legacy program were observed. ZZ Boo and Delta Lib have not been observed since 2016. Both of these stars have periods that are close to an even multiple number of days and are sometimes not observable from a

given longitude for a year or more.

Last year I reported that V342 Aql had not been observed for a few years. In 2018 a single time of minimum was observed. This TOM indicates that another significant change of period has occurred, as shown by the O-C plot below. More observations are needed to follow up on this. This star has a three hour totality so a long observing run is required to obtain a time of minimum.



An O-C plot for V342 Aql from 1935 through 2018. The circled points are times of minima observed with CCD. The rest were observed visually or photographically.

CYCLE

An eclipse of Zeta Aur will occur in the fall of 2019. These eclipses occur every 2.7 years. The table below lists a description of the components.

Component	Α	В
Spectral Type	K4	B7
Diameter (Suns)	200	5
Luminosity	2100	400

As the B7 type star passes behind the K4 type star, most of the light loss will be in the short wavelengths. Because of the size difference, the partial phases of the eclipse are very short, only a couple of days for each. The important dates are: October 24 – 28 for the first and second contacts and December 1 – 5 for the third and fourth contacts. The system will be in total eclipse for the entire month of November. This is a bright star but is a perfect target for those who are observing Epsilon Aur since it is in the same field.



Three eclipses of Zeta Aur observed over the past ten years plotted to phase. Because the loss of light is due to a B7 type star being eclipsed by a much brighter K4 type star, the eclipse depth is negligible in the I band but is a half magnitude deep in the B band. No U band data were available but the eclipse is about 1.5 magnitudes deep in that band.

Exoplanets

Section Leader: Dennis Conti

This past fiscal year (October 1, 2017 – September 30, 2018) has been a busy one for the AAVSO in its support of exoplanet research. This was particularly true in two areas: (1) completion of the AAVSO Exoplanet Database, and (2) the AAVSO's support of TESS (Transiting Exoplanet Survey Satellite) - NASA's latest exoplanet discovery mission.

Similar to its legacy of archiving data on variable stars, the AAVSO, through its Exoplanet Database, can now offer the professional community with an archive of exoplanet timeseries observations by amateur astronomers. Such information can be used to help refine the ephemerides of known exoplanets, as well as to determine Transit Timing Variations (TTVs) that may be indicative of additional companion planets. Exoplanet observations can be submitted through new WebObs forms, and archived data can then later be downloaded by exoplanet researchers.

The most consequential exoplanet event this past year was the launch of TESS. After a successful launch on April 18, 2018, and commissioning, TESS began its science operation. For the first year of operation, TESS will conduct an all-sky survey of the Southern Ecliptic Hemisphere, followed by a year-long survey of the Northern Ecliptic Hemisphere.

2. The Year in Review

Even though it is a space-based telescope, a transit detection by TESS needs to be confirmed by follow-up ground observations since the light from many stars could be blended together in TESS' large photometric aperture. Thus, the TESS science team has established a rigorous pipeline process for moving such a detection to a "confirmed" status. The first step in this pipeline consists of both professional and amateur ground-based observers conducting observations to determine if the TESS detection represents a true exoplanet transit or is a "false positive." A common example of such a false positive is the result of the drop in light of an eclipsing binary near the target star that mimics an exoplanet transit when the light of both are blended together in the TESS aperture The subgroup that is responsible for this determination is the TESS Follow Up Program, Subgroup 1 (SG1).

To facilitate participation in SG1 by its members, the AAVSO established a qualification program whereby members who successfully completed this program would be recommended to the TESS SG1 lead as a TESS follow-up observer. So far, over a dozen AAVSO members have become certified TESS SG1 observers through this qualification program.

In addition to establishing this qualification program, the AAVSO Section Lead developed an extensive guide for conducting and submitting SG1 observations that has now been adopted by the TESS SG1 lead for use by other participants in the SG1 program. Also, the AAVSO Section Chair developed an extension to AstrolmageJ for automatically analyzing the existence of and possibility of near-by eclipsing binaries being the source of TESS false positive transit detections. This feature is now also being used by participants in other exoplanet surveys such as KELT.

During this past year, the AAVSO continued to provide its members with training on exoplanet observing through two CHOICE course offerings. Over 120 AAVSO members have now completed the Exoplanet Observing course since its inception in February 2017.

Also, during this period, updates were also made to "A Practical Guide to Exoplanet Observing." This document is used as material for the exoplanet CHOICE courses and is also being used as a stand-alone document for use by both amateur astronomers and educational institutions interested in exoplanet observing.

Finally, during this past year the Exoplanet Section Leader conducted a number of outreach activities. These included:

1. Presentation at NEAIC 2018: "Amateur Astronomer Participation in the TESS Exoplanet Mission."

- 2. Presentation at NEAF 2018: "The TESS Exoplanet Mission and Amateur Astronomer Participation."
- 3. Presentation at SAS 2018: "Exoplanet False Positive Detection with Sub-meter Telescopes."
- 4. Presentation on amateur astronomer participation in TESS as part of the Johns Hopkins University Applied Physics Lab (APL) Colloquium Series.

Long Period Variable (LPV)

Section Leaders: Andrew Pearce and Frank Schorr

The AAVSO LPV section reached a steady state in 2018 after years of rejuvenation in 2016 and 2017.

The LPV Section web page continues to be fully functioning and can be found at https://www.aavso.org/aavso-long-period-variable-section.

A number of activities were conducted in FY 2017/2018 which included:

- The LPV of the month page was continued successfully throughout the year and was widely publicised in the LPV Forum and in the monthly AAVSO Communications. Brief descriptions and preliminary analysis was performed on the following stars in the last 12 months: R, Aql, R Cnc, T Cam, Chi Cyg, Z Tau, V Boo, S Cas, T UMi, U UMi, R Nor, SS Vir, R LMi and S Aur. Thanks to Frank Schorr and Alfredo Glez-Herrera for contributing articles throughout the year. The Section reiies on contributions from members and observers to keep this series going. Unfortunately we did not receive a lot of support in this regard in 2018 and therefore the frequency of these articles will reduce to quarterly in 2019.
- An interesting article published in the January 2018 AAVSO Newsletter was an analysis of LPV observations over the past 10 years. This was primarily done to ascertain the current observing levels for LPV's across the AAVSO both visually and via CCD as there were concerns that the coverage of LPV's had diminished in recent times. The pleasing aspect arising out of the analysis was that the number of visual observations and the coverage of the LPV Legacy Program stars has remained essentially constant over the last 10 years. Furthermore the number of non-visual observations has varied more than for visual observations on an annual basis and the coverage of LPV Program Legacy stars is significantly less than for visual observations. A future focus area for the Section is concerning another outcome from the analysis which found that for both types of observations a significant fraction (40%) of the stars observed have less than 10 observations per year. There may be an argument to encourage observers to stop following these poorly observed stars and focus more on others

2. The Year in Review

- which have more well defined light curves or where gaps can be plugged in these light curves.
- Stan Walker (previous Director of Variable Stars South) prepared an interesting article on the changing face of LPV observing. Stan focused on the relationship between visual and CCD/PEP observations in the context of how each type of observing technique can contribute to furthering our knowledge of LPV's. The article is contained on the LPV Section web page.
- Good coverage has been maintained on R Aqr throughout the year supporting Chandra and HST observations of this star as requested by Dr. Margarita Karovska (Harvard-Smithsonian Center for Astrophysics) in AAVSO Alert Notices 535, 589 and 600. A total of 939 observations were made in FY 2017/2018. This is a multi-year campaign leading up to the predicted date of the next eclipse of the system in 2022.
- As outlined in AAVSO Alert Notice 645, Michael Stroh (Ph.D. candidate, University of New Mexico) requested AAVSO assistance in monitoring 19 LPVs in support of a VLA campaign studying SiO masers. This campaign runs until October 2019. Nearly 1,400 observations were made from the start of the campaign in July 2018 until end of FY 2017/2018. Observers are encouraged to maintain a good coverage of these stars throughout 2019.
- All LPV Legacy, LPV Legacy South and Percy List stars are included in the AAVSO
 Target Tool which will allow all observers to effectively plan an LPV observing
 program suited to the particular observer.
- Links to recent scientific papers regarding LPVs where AAVSO observations have been used or referenced continue to be updated on a quarterly basis.
- Frank Schorr continues to maintain the LPV Hump page which contains a lot of data and discussion on these interesting longer period Miras which show a hump or double maximum in their light curve.

The primary goals of the LPV section are: to facilitate the long-term observation, both visually and electronically, of the Legacy LPVs in the program; and to promote other scientifically significant LPV targets for observers to follow. We are particularly interested in encouraging and guiding visual observers to include LPVs in their target selection and in building their own observing programs. As with all Sections, the LPV section requires interest from the larger observer community. The AAVSO encourages both LPV observers and users of AAVSO LPV data in their research to get involved with the AAVSO LPV section.

Photoelectric Photometry (PEP)

Section Leader: Tom Calderwood

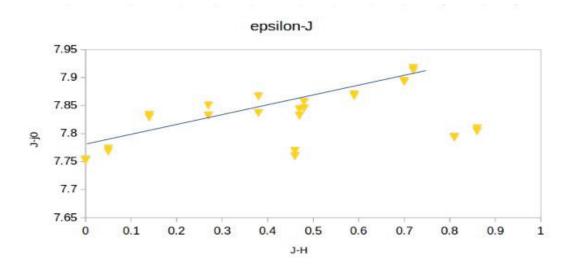
A Near-IR Puzzle – Part I

The PEP group is still working to get the Optec SSP4 near-infrared photometers back into regular use. One challenge is to establish transformation coefficients so that everyone's photometry can be interpreted in a standard system. In the CCD world, B and V transformations are usually established by imaging a "standard" open cluster having stars with a wide range of B-V values. Ideally, the magnitudes you measure will exactly match the standard magnitudes, but the best you can hope for is that your magnitudes exhibit a deviation from standard that is linear in B-V. The transformation value for a band is found by plotting, for each star, the deviation of your extinction-corrected instrumental magnitude from the standard magnitude versus the standard B-V. In photometric lingo, this is V-v0 vs. B-V, (see JH example, below). A least-squares fit is made to the points, and the slope of the line gives the transformation.

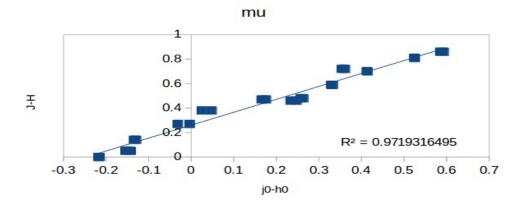
Cluster stars are generally too faint for PEP observers working with photodiode photometers. We use a closely-spaced pair of bright red and blue stars. Our calibration is essentially the same process, except that we "fit" a line to just two stars at opposite ends of the B-V range (reliable magnitudes are essential). There are several suitable pairs in the northern hemisphere, and it is best to establish calibration with more than one pair.

In the near-IR (J and H), there are no usable pairs of reddish/bluish stars. Reliable, bright JH magnitudes are hard to come by, so we need to measure multiple stars in the hope of averaging out the errors. Since we cannot use clusters, we are forced to make an "all-sky" calibration, picking out targets over a considerable range of altitude and azimuth. Below is an example calibration run on thirteen stars, each measured twice.

2. The Year in Review



By tossing out three outlier stars at J-H= 0.46, 0.81, and 0.86, one gets a tolerably good linear fit for the rest. But why are the Terrible Trio so discordant? Are they faint and thus hard for the scope to pick up reliably, or are their magnitudes perhaps not particularly well established? The 0.46 star is rather dim, but the others are plenty bright, and all three are among the more-reliable magnitudes in the UK Infra-red Telescope catalog. Very strange—and this observation run is not a fluke: other calibrations had problems with the same three stars. Further, if you think the epsilon-J plot looks squirrely, the epsilon-H was simply dreadful. The enormous puzzle, however, is μ : the transformation for the J-H color index:



The μ plot is set up differently from the epsilons (J-H versus j0-h0), but the basic idea is the same. Here, the correlation, using all thirteen stars, is excellent. How is this possible given the problems with the epsilons? To be continued....

Short Period Pulsator

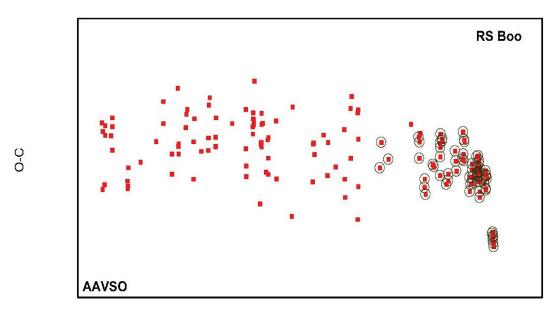
Section Leader: Gerard Samolyk

Once again, all of the stars in the AAVSO legacy program were observed. In 2018, Six of these stars that exhibit the Blazhko effect were heavily observed, SW And, AC And, RS Boo, XZ Cyg, DM Cyg, and AR Her. Most observers only observe these stars near the time of maximum, however observations at all phases including minimum and maximum are important when analyzing Blazhko stars because the brightness and phase at both of these extremes vary during the Blazhko cycle.

A paper containing 375 times of maxima of 86 stars was published in the JAAVSO. This paper contained the reduction of data sent to the section chair by seven observers in 2018. Any observer who would like to contribute data to these papers in the future should upload their observations to the AID and send a copy to gsamolyk@wi.rr.com

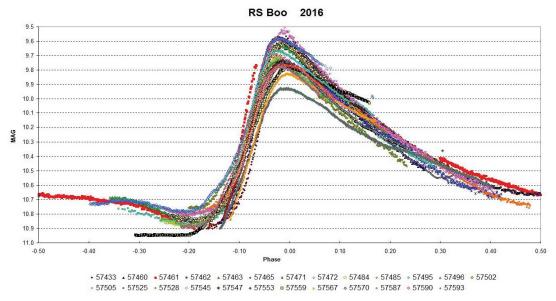
Two more papers containing times of maxima measured from visual observations of RR Lyr stars made by the RR Lyr committee since the 1960's have been published in the JAAVSO. This project is on track to be completed in 2019.

In 2018, a large deviation in the times of maxima for RS Boo was noted (as shown in the O-C plot below). This occurred after the period had been relatively stable for over a half century. This star also exhibits a Blazhko effect, as shown in the phase plot below, we should make this star a priority over the next few years. First, to verify the current period, and second, to determine if the Blazhko period has also changed.



CYCLE

O-C plot for times of maxima of RS Boo using the light elements from 4th edition of the GCVS. Maxima that were observed using CCD's are circled. The vertical scatter is due to the Blazhko cycle.



Observations of RS Boo plotted by phase with each night's observations are color-coded. The variations in magnitude at maximum and minimum as well as phase shifts are evident.

Solar

Section Leader and SID Group Leader: Rodney Howe

Sunspot Group Leader: Kim Hay

Sun Spot Group

The sun is now approaching the quiet time and is becoming somewhat less active, although with ups and downs on a daily basis, as we go into the solar minimum. Kim Hay from Yarker, Ontario, Canada, has done an excellent job of collecting, cleaning, and creating the monthly American Relative numbers for the *Solar Bulletin*. There was a total of 83 observers who contributed 12,558 observations (September 2017 – September 2018). Their efforts should be applauded as they continue to monitor our nearest star. We also have many awards to be given for our sunspot observers based on past certificates, and running numbers for FY 2018 (see table below).

Sunspot Awards 2018

Obs Code	Name	2018 Award
ERB	Bob Eramia	100
DMIB	Michel Deconinck	100
FTAA	Tadeusz Figiel	500
BDDA	Diego Bastiani	500
ONJ	John O'Neill	500
JGE	Gerardo Jimenez Lopez	500
JDAC	David Jackson	500
MWU	Walter Maluf	500
MJAF	Juan Antonio Moreno Quesada	1000
SONA	Andries Son	1000
MUDG	George Mudry	1000
DJOB	Jorge del Rosario	1000
CVJ	Jose Carvajal	1500
MARE	Enrico Mariani	1500
AAX	Alexandre Amorim	1500
KROL	Larry Krozel	2000
HOWR	Rodney Howe	2000

Table continued on next page

2. The Year in Review

Obs Code	Name	2018 Award
VIDD	Dan Vidican	2000
ASA	Salvador Aguirre	2000
AJV	Javier Alonso	2500
BARH	Howard Barnes	2500
LEVM	Monty Leventhal	3500
URBP	Piotr Urbanski	4000
VARG	A. Gonzalo Vargas	4500
DUBF	Franky Dubois	4500
KNJS	James & Shirley Knight	4500
WILW	William M. Wilson	4500
FUJK	K. Fujimori	4500
TESD	David Teske	4500
BROB	Robert Brown	4500
СКВ	Brian Cudnik	5000
ARAG	Gema Araujo	5500
CHAG	German Morales Chavez	6000

SID Group

For the last 12 months, overall SID submissions have been stable. Our observer ranks have remained consistent with between 14 and 17 submissions each month. There was a total of 21 observers submitting reports and a total of 140 reports were sent in for fiscal year 2018. Thanks to all VLF observers for their efforts in monitoring, data analysis and report generation.

Two observers are eligible for the traditional award for over 40 observations this year.

SID Awards 2018

Obs	Name	Level
A-138	Salvador Aguirre	50 reports
A-146	Ken Menzies	50 reports

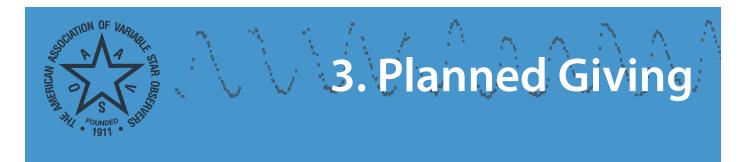
Young Stellar Object (YSO)

Section Leader: Michael Poxon

Once again it is good to see that more observers are coming on board, and with data from a variety of wavebands – not just us visual folks! In this connection the forum announcements and Alert Notices have proved very useful in foregrounding certain objects that need attention. As a very rough-and-ready experiment I called up a few sample light curves and observer details from the LCG. All data below were taken from a 600-day timespan.

Star	No. of Observations	No. of Observers
RW Aur	1583	81
CQ Tau	299	21
V730 Cep	490	16
V1117 Her	872	32
RR Tau	1313	39
UX Ori	432	24
SU Aur	744	11
VX Cas	301	16
YZ Cep	266	10

These figures demonstrate several facts: stars that have been on the AAVSO program for a long time (RW Aur and RR Tau for example) have the greatest number of observations as well as the largest number of separate observers. Stars that have exhibited interesting recent behavior (RW Aur and V1117 Her) also have a high number of observations (quite rightly!), but what surprises me is that the brighter objects in the list such as CQ Tau, YZ Cep and and SU Aur have low numbers of both observations and observers! Bear in mind that CQ Tau is even visible with decent binoculars when bright, and is one of the most active stars on the list. YZ Cep is next to a naked-eye star, and SU Aur is next to another bright YSO. These stars should be on the observing lists of those with small telescopes! Also, rather bizarrely, circumpolar stars like YZ Cep and VX Cas likewise seem to be relatively neglected. I do understand that while there may be some sort of cachet (whether conscious or not) in observing the fainter objects, observers – especially those with very modest equipment – need to follow as many objects as possible. There is something for everyone here, so the small scope owners should not feel left out!



Planned Giving

Charitable contributions to the American Association of Variable Star Observers can have benefits that last a lifetime - and beyond. A bequest or life-income gift that includes the Association will support variable star research and education for generations to come. Your legacy can be made in a variety of ways that can have dramatic tax and financial rewards for donors.

The AAVSO is recognized by the IRS and the state of Massachusetts as a 501c3 nonprofit scientific and educational organization. Gifts of all denominations are welcomed, and may include cash, securities, and other gifts. Unrestricted contributions may be made in any amount, and are tax-deductible to the extent allowed by the law.

The following information is intended only to show some ways by which you can make a charitable gift or bequest and thereby minimize federal tax liabilities. These examples are of a general nature only and should not be applied to your specific situation without first consulting your attorney or financial adviser.

A Gift of Appreciated Stock

Appreciated Stock makes an excellent charitable gift. Under current tax laws, when an appreciated asset (such as stock) is sold, a capital gains tax is due. By making a charitable gift of the appreciated stock, you can avoid or delay the capital gains tax. You can also take an immediate income tax deduction for the current fair market value of the stock, no matter what was originally paid for it.

To take a deduction for gifts of appreciated stock at their current value, you must have owned the stock for more than 12 months. Such gifts are deductible up to 30 percent of your adjusted gross income in the year of the gift. Any unused deduction amounts may be used in as many as five subsequent tax years.

A Bequest Through Your Will

One of the most simple and popular ways to make a gift that will live after you is to give through your will. You can make a gift bequest to benefit the AAVSO by providing a dollar amount, specific property, percentage of your estate, or the remainder (what's left). Such a designation can reduce your estate taxes. In many cases, a simple change to your will can add this bequest and does not require rewriting your most recent will.

Create a Charitable Remainder Trust

Donors and spouses can benefit from lifelong payments from such a trust. The donor selects the rate of return from these income arrangements and also chooses a fixed or fluctuating annual payment to be made to the designated parties as long as they live. Capital gains tax may be completely bypassed and you will receive a tax deduction based on the age of the income recipient and the rate of return chosen.

Establish a Charitable Lead Trust

In a charitable lead trust, assets (generally cash or securities) are transferred to a trust that pays income from the fund to your favorite charity or charities for the number of years you select. At the end of the designated time period, the trust terminates and the assets are given back to the person you name. This trust helps to lower estate and gift taxes that would otherwise be due on the assets. This option is especially attractive if you want to leave your children or grandchildren assets in the future, but not immediately.

Design a Gift Annuity

In exchange for a gift of cash, stock, or securities, the AAVSO will pay you, you and your survivor, or another person named by you, a guaranteed income for life. In addition, you receive a substantial income tax deduction in the year of the gift, and part of the annual payment is non-taxable. Your annuity payment and tax deduction are based on the age and income of the recipient.

Consider a Deferred Gift Annuity

A deferred gift annuity is similar to a gift annuity except that payments begin for you at a future date of your choice, such as your retirement. Your tax deduction and the annual rate of return on your annuity increase the longer you wait to start payments. This is an excellent retirement planning method to implement during prime income producing years that will benefit you later.

3. Planned Giving

Use Your Life Insurance

Insurance is another simple way to make a substantial future gift at a level that would not be possible at the same level in cash. Name the AAVSO the owner and beneficiary to receive the proceeds of an existing life insurance policy. You will receive a tax deduction for approximately the cash surrender value, thereby reducing your tax liability in the year of the gift.

An alternative is to purchase a new life insurance policy naming this organization as owner and beneficiary. With this option, you receive an income tax deduction for each premium made and make possible a major gift to your favorite charity with a modest annual payment (or one-time premium payment).

Look to Retirement Accounts for Gift Opportunities

Account funds (IRAs or company plans) beyond the comfortable support of yourself or loved one may be given (like life insurance proceeds) to the AAVSO by proper beneficiary designation. Large pension plan assets can be subject to double or triple taxation (federal estate, federal income, and state death and state income tax) that virtually eliminates the benefits to heirs if tax-wise alternative planning is not arranged.

Plan a Gift of Real Estate

For some people, a gift of land, primary residence, or vacation home is a preferred way to make a gift. You will receive a tax deduction for the full fair market value, avoid all capital gains tax, and remove this asset from future estate taxes. One option is to give real estate while you retain a life tenancy. This provides a substantial income tax deduction by giving (deeding) your home or farm to your favorite charity now. You continue to live there, maintain the property as usual, and even receive any income it generates. At your death the property will be sold and the proceeds will support the mission of the AAVSO.

We are very grateful to all our donors, who are listed in Appendix 4.



Appendix 1. Observer Awards

AAVSO Observer Awards (presented at the 107th Annual Meeting in Flagstaff, Arizona, November 17, 2018)

Award/recipient	Affiliation**	Country	Interval	Total
OVER 125,000 VISUAL OBS	SERVATIONS*			
Frank Vohla	02	Germany	1989-2018	131,392
OVER 100,000 VISUAL OBS	SERVATIONS*			
Andrew Pearce		Australia	1989-2018	101,834
Sandor Papp	03	Hungary	1983-2018	101,613
OVER 75,000 VISUAL OBSE	ERVATIONS*			
Kenneth J. Medway		Great Britain	1973-2018	75,078
OVER 25,000 VISUAL OBSE	ERVATIONS*			
Alexandr S. Maidik		Ukraine	1998-2018	25,371
OVER 10,000 VISUAL OBSE	ERVATIONS*			
Antonio Padilla Filho		Brazil	1987-2018	10,336
OVER 5,000 VISUAL OBSER	RVATIONS*			
Peter Molnar	03	Hungary	2003-2018	8,005
Robert F. Swanberg		USA	1962-2018	7,866
Walter J. Kaminski		USA	1959-2018	7,014
Thomas C. Bretl		USA	1975-2018	6,515
Albert G. Dill		USA	1993-2018	6,175
Ivan Sergey		Belarus	2004-2018	6,087
Sherrill Shaffer		USA	2008-2018	6,015
Geert Hoogeveen	04	Netherlands	1975-2018	5,600
Laszlo Kocsmaros		Rep. of Serbia	2009-2018	5,231
Arthur Sturm	02	Germany	1996-2018	5,096
OVER 1,000 VISUAL OBSER	RVATIONS*			
Javier Alonso	15	Spain	2004-2018	4,941
			continued	on next page

Appendix 1. Observer Awards

Award/recipient	Affiliation**	Country	Interval	Total
Mark A. Bradbury		USA	1992-2018	4,566
Edwin Van Dijk	04	Germany	1995-2018	3,379
Vincent R. Coulehan		USA	1972-2018	3,267
Andrew Beaton	27	Canada	2002-2018	3,258
Lance E. Shaw		USA	1996-2018	3,067
Jerome A. Hudson		USA	1960-2018	2,359
Ivan Vincze	03	Hungary	1990-2018	2,260
Paulo S. Reis Fernandes	13	Brazil	1997-2018	1,647
Tony Vale	20	Great Britain	2015-2018	1,316
Michel Deconinck		France	2014-2018	1,300
Bodo Wichert		Germany	2015-2018	1,255
Steven L. Koontz		USA	2016-2018	1,099
Igor Yatsenkov		Russia	2014-2018	1,043
Dominique Proust	01	France	1980-2018	1,037
OVER 100 VISUAL OBSERVA	ATIONS*			
Hernan De Angelis		Sweden	2017-2018	949
Frank Krafka		USA	2001-2018	687
Keith Geary		Ireland	2002-2018	679
Andy Howell		USA	1964-2018	665
Richard Roberts		USA	2017-2018	660
Tamas Jakabfi	03	Hungary	2003-2018	469
Marc Eichenberger		Switzerland	2003-2018	409
Gabriel Murawski		Poland	2012-2018	370
Andreas Kammerer	02	Germany	2001-2018	358
Raymond B. Howard		USA	2017-2018	308
Katarina Holmquist	19	Sweden	2014-2018	298
Benny Colyn	05	Belgium	1998-2018	236
Zoltan Magyarics	03	Austria	2017-2018	211
Irina Sagaybins'ka		Ukraine	2018-2018	209
Richard A. Loslo		USA	2017-2018	200
Marcin Figuła		Poland	2018-2018	198
Mike Stone		USA	2017-2018	193
Elena Odintsova		Ukraine	2018-2018	180
Charles F. Trefzger		Switzerland	1990-2018	163
Roberto Valdomiro Silva	а	Brazil	2017-2018	162

Award/recipient	Affiliation**	Country	Interval	Total
Charles E. Stewart		USA	2018-2018	161
Jason J. Rivet		USA	2016-2018	137
William H. Carini		USA	1985-2018	127
Cledison Marcos da Si	lva	Brazil	2017-2018	126
Preston Pendergraft		USA	2016-2018	125
Jean Pierre Dechoz		France	2016-2018	125
Arturo Purroy		Spain	2013-2018	112
Szabolcs Cziniel	03	Hungary	1992-2018	110
Becky Ramotowski		USA	2001-2018	108
Fabio Fujiwara	13	Brazil	2012-2018	106
Jonathan Poppele		USA	2016-2018	106
OVER 2.0 MILLION CCD O	RSFRVATIONS*			
Franz-Josef Hambsch	05	Belgium	2002-2018	2,062,698
		3		
OVER 800,000 CCD OBSE	RVATIONS*			
Shawn Dvorak		USA	1981-2018	851,595
OVER 700,000 CCD OBSE	RVATIONS*			
Teofilo Arranz	IVAIIONS	Spain	2005-2018	794,631
7.07.07.07.07.0		Spa	2003 2010	7 7 1,00 1
OVER 600,000 CCD OBSE	RVATIONS*			
Gerard Samolyk		USA	1975-2018	632,611
OVER 500,000 CCD OBSE	RVATIONS*	LICA	2016 2010	E44 E22
Geoffrey Stone	0.5	USA	2016-2018	544,533
Tonny Vanmunster	05	Belgium	1976-2018	515,413
OVER 400,000 CCD OBSE	RVATIONS*			
David Cejudo Fernanc		Spain	2010-2018	442,782
Gordon Myers		USA	2007-2018	409,852
Lewis (Lew) Cook		USA	1974-2018	408,063
. ,				
OVER 300,000 CCD OBSE	RVATIONS*			
Joseph H. Ulowetz		USA	2010-2018	303,288

Appendix 1. Observer Awards

Observer Awards, cont.				
Award/recipient	Affiliation**	Country	Interval	Total
OVER 100,000 CCD OBSER	/ATIONS*			
James A. Boardman		USA	2012-2018	105,189
Tamas Tordai	03	Hungary	1986-2018	104,615
Douglas E. Barrett		France	2011-2018	101,425
Timo J. Kantola		Finland	2009-2018	101,012
Yenal Ogmen		North Cyprus	2004-2018	100,969
Maarten Vanleenhove	05	Belgium	2014-2018	100,127
OVER 50,000 CCD OBSERVA	ATIONS*			
Stephen M. Brincat		Malta	1984-2018	56,530
Laurent Corp		France	1997-2018	54,352
John W. Rock		Great Britain	2012-2018	62,879
OVER 10,000 CCD OBSERVA	ATIONS*			
Pablo Lewin		USA	2017-2018	24,573
Sjoerd Dufoer	05	Belgium	2001-2018	22,315
Jim Seargeant		USĂ	2015-2018	19,538
Marco Ciocca		USA	2011-2018	11,539
Ivaldo Cervini		Switzerland	2017-2018	11,051
Tadeusz Smela		Poland	2014-2018	10,669
Tom Polakis		USA	2017-2018	10,367
Robert B. Jenkins	34	Australia	2013-2018	10,273
OVER 1,000 CCD OBSERVA	TIONS*			
Matthew F. Knote		USA	2018-2018	6,582
Kristie B. Caplikas		Australia	2018-2018	5,546
Dan Crowson		USA	2017-2018	5,446
Mario A. Morales Aimar		Spain	2015-2018	4,849
Gabriel Murawski		Poland	2012-2018	4,525
John F. Briol		USA	2012-2018	4,226
Heinz-Bernd Eggensteir	า	Germany	2009-2018	3,893
Richard E. Schmidt		USA	2018-2018	3,034
Joseph P. Seitz		USA	2017-2018	2,558
Claudio Lopresti	18	Italy	2017-2018	2,420
Rafael Gonzalez Farfan		Spain	2014-2018	1,959

Observer Awards, cont.

Award/recipient	Affiliation**	Country	Interval	Total
Jean-Francois Gout		USA	2016-2018	1,741
Francisco Sevilla Lobat	:0	Spain	1996-2018	1,594
Jeff W. Robertson		USA	2013-2018	1,583
William S. Rea	29	New Zealand	2015-2018	1,581
Jean-Bruno Desrosiers		Canada	2014-2018	1,529
Klaus Wenzel	02	Wenzel	2004-2018	1,401
Andrew Pearce		Australia	1990-2018	1,330
Matthew S. Bundas		USA	2017-2018	1,264
Rob Solomon		Australia	2014-2018	1,236
Stanley Shadick	27	Canada	2011-2018	1,204
Charles Galdies		Malta	2017-2018	1,183
John C. Ruthroff		USA	2008-2018	1,165
Huei Sears		USA	2017-2018	1,091
Dmitry Matsnev		Russia	1991-2018	1,079
Erik J. Schwendeman		USA	2016-2018	1,062
Mervyn Millward		Australia	2011-2018	1,055
OVER 10,000 PEP OBSERV	ATIONS*			
Gerald Persha		USA	2013-2018	11,052
OVER 10,000 DSLR OBSER	VATIONS*			
Aniruddh N. Deshpano	de	India	2013-2018	10,134
OVER 1,000 DSLR OBSERV	'ATIONS*			
Wolfgang Vollmann		Austria	1976-2018	2,475
Erik Wischnewski	02	Germany	2014-2018	2,464
Richard Biernikowicz		Poland	2011-2018	2,427
Nicola Montecchiari	18	Italy	1990-2018	2,289
Ivan Sergey		Belarus	2003-2018	1,262
David L. Blane	10	South Africa	1978-2018	1,095
OVER 500 DSLR OBSERVA	TIONS*			
Maksym Pyatnytskyy		Ukraine	2017-2018	978
Stephen C. Schiff		USA	1960-2018	834
Søren C. Toft		Switzerland	2016-2018	579

Appendix 1. Observer Awards

Observer Awards, cont.

Award/recipient	Affiliation**	Country	Interval	Total
OVER 100 DSLR OBSERVA	TIONS*			
Gregory Conrad		USA	2017-2018	499
Ray E. Tomlin		USA	2006-2018	498
Attila Nemes		Hungary	2006-2018	444
Francisco Jablonski		Brazil	2018-2018	391
John A. Pickett		USA	1988-2018	322
Robert K. Buchheim		USA	2005-2018	310
Dietmar Augart	02	Germany	2002-2018	243
Yuri Solomonov		Russia	2017-2018	234
David Mohrbacher		USA	1980-2018	209
Piotr Ligeza		Poland	2018-2018	165
Sampsa Lahtinen	17	Finland	2004-2018	152
Adam Derdzikowski		Poland	2003-2018	150
Bogdan Kubiak		Poland	2009-2018	144
Michael G. Miller		USA	2010-2018	143
Mark G. Blackford	29	Australia	2011-2018	117
David Rodriguez		Spain	2016-2018	114
David J. Benn	34	Australia	2009-2018	112
Christian Riou		France	2012-2018	104

^{*} Years include total AAVSO observing interval (not only PEP/CCD observing). Total includes only visual or PEP/CCD observations, depending on award.

A number preceding a name indicates observer is also affiliated with the group below:

- 01 Association Française des Observateurs d'Étoiles Variables (AFOEV)
- 02 Bundesdeutsche Arbeitsgemeinschaft für Veränderliche Sterne e.V.(Germany) (BAV)
- 03 Magyar Csillagászati Egyesület, Változócsillag Szakcsoport (Hungary) (MCSE)
- 04 Koninklijke Nederlandse Vereniging Voor Weer-en Sterrenkunde, Werkgroep Veranderlijke Sterren (Netherlands) (KNVWS)
- 05 Vereniging Voor Sterrenkunde, Werkgroep Veranderlijke Sterren (Belgium) (VVS)
- 10 Astronomical Society of Southern Africa, Variable Star Section (ASSA)
- 13 Rede de Astronomia Observacional (Brazil) (REA)
- 15 Agrupacion Astronomica de Sabadell (Spain) (AASABADELL)
- 17 URSA Astronomical Association, Variable Star Section (Finland) (URSA)
- 18 Unione Astrofili Italiani (UAI)
- 19 Svensk Amator Astronomisk Forening, variabelsektionen (SAAF)
- 20 British Astronomical Association, Variable Star Section (BAA-VSS)
- 27 Royal Astronomical Society of Canada (RASC)
- 29 Variable Stars South (VSS)
- 34 Astronomical Society of South Australia (ASSAU)

AAVSO Solar Observer Awards (presented at the 107th Annual Meeting in Flagstaff, Arizona, November 17, 2018)

Sunspot Award (100 observations)

Michel Deconinck

Bob Eramia

Sunspot Award (500 observations)

Diego Bastiani Tadeusz Figiel

David Jackson

Gerardo Jimenez Lopez

Walter Maluf John O'Neill

Sunspot Award (1,000 observations)

Jorge del Rosario

Juan Antonio Moreno Ouesada

George Mudry Andries Son

Sunspot Award (1,500 observations)

Alexandre Amorim

Jose Carvajal Enrico Mariani

Sunspot Award (2,000 observations)

Salvador Aguirre Rodney Howe Larry Krozel

Dan Vidican

Sunspot Award (2,500 observations)

Javier Alo

Howard Barnes

Sunspot Award (3,500 observations)

Monty Leventhal

Sunspot Award (4,000 observations)

Piotr Urbanski

Sunspot Award (4,500 observations)

Franky Dubois

K. Fujimori

James & Shirley Knight

A. Gonzalo Vargas

William M. Wilson

Sunspot Award (5,000 observations)

Robert Brown

Brian Cudnik

David Teske

Sunspot Award (5,500 observations)

Gema Araujo

Sunspot Award (6,000 observations)

German Morales Chavez

SID Award (40 months of reports)

Salvador Aguirre

Ken Menzies



Table 1. AAVSO Observer Totals 2017–2018 by Country.*

Country	No. Observers	No. Obs.	Country	No. Observers	No. Obs.	Country	No. Observers	No. Obs.
Argentina	10	695	France	33	33953	Romania	8	1085
Austria	7	2841	United Kingdom	68	185110	Serbia	1	430
Australia	30	64976	Greece	8	12035	Spain	52	344258
Bangladesh	1	4	Croatia	1	44	Russian Federation	16	1045
Belgium	21	382484	Hungary	33	67478	Sweden	10	8946
Bulgaria	3	1672	Ireland	3	102	Singapore	10	21
Bermuda	1	292	India	6	6017	Slovenia	1	22
Bolivia	1	52	Italy	33	15067	Slovakia	6	2676
Brazil	22	4707	Japan	3	1061	Turkey	2	11
Belarus	3	520	Kazakhstan	1	123	Taiwan	1	179
Canada	42	33136	Luxembourg	1	7	Ukraine	16	16451
Switzerland	6	9752	Latvia	1	8	United States	272	1115069
China	1	1840	Morocco	1	53	Uruguay	1	111300
Cyprus	2	16533	Malta	2	34109	Venezuela	1	46
Czech Republic	1	1	Mexico	1	139	South Africa	4	747
Germany	40	26903	Netherlands	7	2391	Joddii Airica		, , ,
Denmark	8	1773	Norway	3	184	Total countries:	56	
Ecuador	1	17	New Zealand	8	16653	Total observers:	846	
Estonia	1	69	Philippines	1	719	Total observations:	2435646	
Spain	52	344258	Poland	26	8547	iotai observations.	2433040	
Finland	10	10312	Portugal	3	2276			

Table 2. AAVSO Observer Totals 2017–2018 USA by State or Territory.

Obs. State Maryland Massachusetts Michigan Minnesota Mississippi Missouri Montana	7 10 10 4 1 5 2	Obs. 2731 52306 6509 3425 1606 6436 32254	South Carolina Texas Utah Vermont Virginia	Observers 2 16 2 1 14	Obs. 155 14784 8 1147
70 Massachusetts 18 Michigan 17 Minnesota 73 Mississippi 59 Missouri 4 Montana	10 10 4 1 5	52306 6509 3425 1606 6436	Texas Utah Vermont Virginia	16 2 1	14784
Michigan Minnesota Mississippi Missouri Montana	10 4 1 5	6509 3425 1606 6436	Texas Utah Vermont Virginia	16 2 1	14784
Minnesota Mississippi Missouri Montana	4 1 5	3425 1606 6436	Utah Vermont Virginia	2	8
73 Mississippi 59 Missouri 4 Montana	1 5	1606 6436	Vermont Virginia	1	-
59 Missouri 4 Montana	_	6436	Virginia	1 14	1147
4 Montana	_			14	
	2	32254		1.7	3384
		32231	Washington	8	496
34 Nebraska	1	24	West Virginia	2	1461
93 Nevada	1	1	Wisconsin	5	126216
13 New Hampshire	4	2796	unknown	3	905
27 New Jersey	3	3452			
84 New Mexico	9	58402	Total states:	46	
30 New York	11	39686	Total observers:	272	
16 North Carolina	6	5666	Total observations:	: 1115069	
55 Ohio	13	3714			
53 Oklahoma	3	418			
52 Oregon	5	27763			
61 Pennsylvania	8	1915			
	1	4668			
5	Oklahoma Oregon	63 Oklahoma 3 62 Oregon 5 61 Pennsylvania 8	63 Oklahoma 3 418 62 Oregon 5 27763 61 Pennsylvania 8 1915	63 Oklahoma 3 418 62 Oregon 5 27763 61 Pennsylvania 8 1915	63 Oklahoma 3 418 62 Oregon 5 27763 61 Pennsylvania 8 1915

^{*}Totals reflect observations made during fiscal 2017–2018 and do not include historical data (data preceding fiscal 2017–2018) submitted during fiscal 2017–2018.

Table 3. AAVSO Observers, 2017–2018.*

				No.					No
ode	Org.		Name	Obs.	Code	Org.		Name	Ob
ΑP		P.	Abbott, Canada	55	BJOI		J.	Black, New Zealand	
AN	02	A.	Abe, Germany	48	BRAC		R.	Black, Oklahoma	4
PGA	20	P.	Abel, United Kingdom	6	BMGA	29	M.	Blackford, Australia	55
GUA		G.	Abramson, Argentina	1	BKL		J.	Blackwell, New Hampshire	4
PIA	18	P.		215	BADB		A.	Blake, Arizona	6
CN	13		Adib, Brazil	120	BVZ		J.	Blanco Gonzalez, Spain	1
SA		S.		139	BLD	10	D.		66
SW	20	S.	· · · · · · · · · · · · · · · · · · ·	899	BWZ		E.		8
FSA	20	F.		696	BJAA		J.	•	2402
CO	20	C.		944	BOH	02		Boehme, Germany	2(
JV	15	J.	Alonso, Spain	504	BSCC	02	S.		48
KV	13		Alton, New Jersey	3429	BHQ	29	T.		48
AX	36		Amorim, Brazil	1924	BRJ	23	J.	•	180
ADA	12			1924	BMF	27		•	100
		Α.	, 3			27		Boschat, Canada	
JN	27	J.	Appleyard, Canada	11	BDLA	0.4		Boulet, Delaware	
RJ		J.	Arnold, Texas	144	BMU	04		Bouma, Netherlands	4
TE		T.		218421	BJMC		J.	5 . 5	
PAA	27	P.	Ashmore, Canada	111	BDG	20		Boyd, United Kingdom	803
ΓDΑ		T.	· · · · · ,	4	BMK			Bradbury, Indiana	:
AUA	32		Audejean, France	152	BMAK		M.	Bradley, Ohio	
DI	02		Augart, Germany	189	BRAF		R.		
RX		R.	Axelsen, Australia	3	BJFA		J.	Brandie, China	18
CJB		C.	Ayres, Nevada	1	BQC	01	J.	Breard, France	
MID		M.	Aznar Carbo, Spain	193	BTB		T.	Bretl, Minnesota	
ΟZ	03	B.	Bago, Hungary	1399	BHA	02	Н.	Bretschneider, Germany	2
IUB		J.	Bagues, Spain	1	BRCM	20	C.	Briden, United Kingdom	
WΥ		W.	Bailey, Arizona	174	BQE	27	E.	Briggs, New York	
MAM		M.	Bajer, Poland	1	BSM		S.	Brincat, Malta	326
JMB		J.	Baker, Michigan	1	BJFB		J.	Briol, Minnesota	329
FO	03	J.	Bakos, Hungary	3547	BSTA		S.		64
DAD		D.		63	BLUA		L.	Brooks, Virginia	1:
GΖ		G.		737	BBM		В.	Brown, Washington	
TAD		T.		1	BNK			Brown, Australia	17.
SBB		S.		15	BLQ	20	L.	Brundle, United Kingdom	134
LOC	18	L.	Barbieri, Italy	95	BANG	20	Α.		11
MAI	10		Barlazzi, Italy	239	BOA	01	Α.		300
SR	18	S.		42	BJRA	O I	J.	Bruton, California	54
PO	10		Barrett, France	11040	BYQ		л. Т.		
ARM	20			407	BISA	16		Bryant, Maryland	
	20		Barrett, United Kingdom			10	l.	Bryukhanov, Belarus	
SAA		S.	, ,,	6	BVIB		V.	•	2
SAB		S.	* **	1	BHU		R.	•	3
3A		В.	•	2926	BNBA			Buchholz, Germany	14
ΝX	27	Α.	•	46	BMSA			Bundas, Michigan	12
QC		A.	. 3	267	BIW	29	N.	Butterworth, Australia	6
ANL		A.	Belotsky, Russian Federation	75	CALC		A.	, ·	1
ZX		A.	Beltran, Bolivia	52	CPU	13	P.	Cacella, Brazil	8
VQ.		R.	Benavides Palencia, Spain	712	CTOA		T.	Calderwood, Oregon	
1S		Н.	Bengtsson, Sweden	266	CFJA		F.	Caleya Salamanca, Spain	
DJB	34	D.	Benn, Australia	118	CCB		C.	Calia, Connecticut	
Υ		T.	Benner, Pennsylvania	445	CCZ		C.	Calis, Turkey	
PAD		P.	Benni, Massachusetts	756	CLUB	36	L.		
В		R.	Berg, Indiana	13	CMN		R.		
EC	27	S.	Bergeron, Canada	580	CMP		R.		12
CA	36	J.	Bernardes, Brazil	63	CFRA		F.	• •	4
/F	50	н. Н.		553	CQP		Α.		1
RIA							K.		
	OΓ	R.	,	1312	CKBB				55
31	05	B.	, 3	393	CMAE	21	Μ.	11 ' /	
COA		C.	1, 3	7	CADA	36	Α.	•	•
XT	80	T.	Bjerkgaard, Norway	182	CJEC		J.	Carels, Belgium	

Table 3. AAVSO Observers, 2017–2018, cont.*

Code	Ora.	Name	No. Obs.	Code	Org.		Name	No. Obs
	Org.		I			_		
CW		W. Carini, New York	-	DJEG DJEF	05	J.	, 3	3
CLQ		L. Cason, South Carolina				J.	Dechoz, France	28
CLAC	0.1	L. Cassignard, France		DMIB DDAA			Deconinck, France	487 18
CJE	01	J. Castellani, France		DLM	01		Dedrickson, Oregon Deldem, France	101
CRAB CDZ		R. Castillo, Spain	<u> </u>	DJEC	ΟI	J.		101:
CIVA		D. Cejudo Fernandez, Spain I. Cervini, Switzerland	20.20	DFR	27	F.	, ,	4
CJOC		J. Chan, Singapore	0.0.	DDE	21		Denisenko, Russian Federation	100
INT		D. Chantiles, California		DAT			Derdzikowski, Poland	280
GF		G. Chaple, Massachusetts	.55	DNO			Deren, Poland	654
CGZ	20	G. Chaplin, United Kingdom		DADA			Deshpande (Ameya), India	1
CAEA	20	A. Chapman, Argentina		DAND			Deshpande (Aniruddh), India	5960
CCY		C. Chiselbrook, Georgia		DJED		J.	•	948
CMF	02	M. Chudy, Germany	323	DPK		P.	Detterline, Pennsylvania	
MAA	02	M. Ciocca, Kentucky	-/ .	DEY		J.		37
CWJA		W. Clark, United Kingdom	3232	DSI			Di Scala, Australia	316
WP.		W. Clarke, Arizona		JNDB			Dias Cavalcante, Brazil	
PE		P. Closas, Spain		DMAC	06		Díaz, Spain	4:
.PP		P. Coker, Colorado		DPA	05		Diepvens, Belgium	39
DK		D. Collins, North Carolina		DIE		S.		
ME	18	E. Colombo, Italy		DLA		A.	Dill, Kansas	9
TIA		T. Colombo, Italy		DMIC	34		Doherty, Australia	3
BY	05	B. Colyn, Belgium		DXAA	15		Domingo Martínez, Spain	232
DSA	20	D. Conner, United Kingdom		DSN		S.	Donnell, Colorado	4
GRD		G. Conrad, New Mexico	1760 I	DROD		R.	Donner, New York	
EMB	01	E. Conseil, France	14	DJUA		J.	Dos Santos, Argentina	93
GIA		G. Conzo, Italy	563 I	DERA		E.	Dose, New Mexico	2508
00		L. Cook, California	66900 I	DDJ			Dowhos, Canada	214
MJA		M. Cook, Canada	3273 I	DRCA	20	R.	Dryden, United Kingdom	112
:K		S. Cook, Arizona	1827 I	DUBF	05	F.	Dubois, Belgium	5956
:WT		W. Cooney, Texas	1295 I	DPV		P.	Dubovsky, Slovakia	1607
COM	10	T. Cooper, South Africa	45 I	DROB	27		Dudley, Ohio	
LZ		L. Corp, France	5016 I	DFS	05	S.	Dufoer, Belgium	2159
MEA		M. Correa, Spain		DMO	01		Dumont, France	73
Al		A. Correia, Portugal	,	DMPA			Durkin, New York	34
NQ		N. Costa, Portugal		DKS		S.	,	144259
OV		V. Coulehan, New York		DJAE		J.	, ,	33
DJA		D. Coulter, Idaho		ELIA		L.	Eager, Florida	
WD		D. Cowall, Maryland		ELYA		L.	,,	13
MY		M. Crook, United Kingdom		ETOA		Τ.	•	69
DAD		D. Crowson, Missouri		EHEA			Eggenstein, Germany	409
DVA	03	D. Csillag, Hungary	· .	EMA			Eichenberger, Switzerland	
SM	03	M. Csukas, Romania		EELA EWRA			Ejarque Gonzalez, Spain	:
KB		B. Cudnik, Texas	525.	EDAB	27		Elias, Argentina	117:
SEB		S. Curry, Oregon	, ,	EPIA	21		English, Canada Ercoli, Italy	5:
CHD	02	C. Cynamon, Maryland		EEY			Erdelyi, California	45:
SZ OCMA	03	S. Cziniel, Hungary C. Da Silva, Brazil		ERW	14	R.	· · · · · · · · · · · · · · · · · · ·	43.
MIA		M. Dadighat, California		FRGA	17	R.	•	185
AH	08	H. Dahle, Norway		FFAD		F.	Feijo, Brazil	21
PHA	00	P. Danthine, Belgium		FJCB	36	J.	• •	19
DAM		A. Darriba Martinez, Spain		FPAA	55	Э. Р.		1.
AJ		J. Davis, Maryland		FRF	03		Fidrich, Hungary	·
)MA		M. Davis, South Carolina		FMAE	03		Figula, Poland	20
HEB		H. De Angelis, Sweden		FMAC			Filipek, Poland	38
ILD		J. De Elías, Spain		FGP	20		Fleming, United Kingdom	40
IAD		J. De Liias, Spaili			03		Fodor, Hungary	
	27	M De long Canada						
DJAD DJX DPP	27	M. De Jong, CanadaP. De Ponthiere, Belgium		FDA FBZ	03	В.		131 42

Table 3. AAVSO Observers, 2017–2018, cont.*

Code	Org.		Name	No. Obs.	Code	Org.		Name	No. Obs
FJQ		J.	Foster, California	993	HBB		В.	Harris, Florida	1614
RL		R.	•	245	HMQ			Harris, Georgia	1.
DU			Fowler, Ohio	16	HWB	20		Harris, United Kingdom	1
ΧJ		J.	Fox, New Mexico	165	HHU	05	Н.	Hautecler, Belgium	
ALD		A.	Free, Washington	4	HAB		R.	Hays, Illinois	49
TKB		T.	Friedli, Switzerland	18	HMH		M.	Heald, New Mexico	23
MG		G.	Fugman, Nebraska	24	HKIC		K.	Heneault, Georgia	
FAB	13	F.	Fujiwara, Brazil	18	HMV			Hessom, California	3.
RTA			Fuller, Texas	10	HNDA			Hewitt, United Kingdom	1
SRA		S.	. 3	369	HXT	20		Heywood, United Kingdom	651
GSTB	20	S.	3 , 3	6	HKEB			Hills, United Kingdom	452
SMQA	20		Gainsford, United Kingdom	422	HDHA			Hinzel, Virginia	26
GCHB GFDB	06		Galdies, Malta	1452 181	HDF HGP	20		Hohman, New York	11.
3FDB 3AA	00	F. P.	Garcia, Spain Garey, Illinois	69	HGUA	20 19		Holahan, Pennsylvania Holmberg, Sweden	115
GJP		г. J.	Garlitz, Oregon	61	HKAB	19		Holmquist, Sweden	23:
GALB		Α.	· . •	153	HOO	04		Hoogeveen, Netherlands	174
GVQ	20	D.		194	HJG	0-1	J.	5	2
GKI			Geary, Ireland	11	HMID			Mir Sakhawat, Bangladesh	_
GMD			Geldorp, Canada	2	HMIA			Hotka, Colorado	17-
GVAA		V.	• *	5	HSP	14		Hovell, New Zealand	444
GDRA		D.		2	HRBA		R.	Howard, California	29
GQR		R.	Gherase, Romania	18	HOA		A.	Howell, Florida	
3MAD		M.	Gibaja, Spain	6	HMAF		M.	Hrivnák, Slovakia	8.
GGU	04	G.	Gilein, Netherlands	747	HJA		J.	•	2
SSEB		S.	Girard, Oklahoma	53	HGAC		G.	Huffman, Alabama	
GRIB		R.		31	HKD	20		Hunt, United Kingdom	41
GATH		Α.	•	4	HUR	20		Hurst, United Kingdom	139
GZN			Glez-Herrera, Spain	4051	HUZ			Huziak, Canada	54
GLG			Gliba, Maryland	130	ILUA		L.	Izzo, Italy	
GFB	31		Goff, California	2197	JFRA		F.	,	39
GPX			Goltz, Australia	36	JPHA	11	Р.	Jackman, Iowa	
GTIA GFDA	27	T. F.	3,	19 1	JJB JMA	11	J.	•	80
GCJ	21	г. J.	Gonzalez, Canada González Carballo, Spain	328	JTP	01	P.	Jacquesson, France Jacquet, France	469
3DJA			Gorney, Arizona	4	JDAA	ΟI	D.		161
GGC			Gotta, Italy	1	JM			James, New Mexico	1923
GJED		J.	Gout, Mississippi	1606	JMAA			Jansson, Sweden	1,723
SMIB			Govedič, Slovenia	22	JDAD			Janzen, Canada	10
GΚΑ		K.		1947	JDAB			Jarkins, Missouri	5:
SPE			Grainger Observatory, New Hampshire	3	JRBA	34		Jenkins, Australia	244
GLAA	01	L.	- · · · · · · · · · · · · · · · · · · ·	940	JSI		S.	Jenner, United Kingdom	4
GRL	80	B.	Granslo, Norway	1	JGE	06	G.	Jiménez López, Spain	5
SANG		A.	Gregory, Minnesota	64	JSJA	20	S.	Johnston, United Kingdom	117
iFS	20	K.	Griffiths, United Kingdom	118	JJI		J.	Jones, Oregon	2754
SALD	03		Grosz, Hungary	18	JMAB			Jones, Indiana	4
GCO			Gualdoni, Italy	1267	JPG		P.	Jordanov, Bulgaria	20
SFRB		F.	Guenther, Maryland	56	JJNA		J.		
ARB		Α.		2	JGRB			Joshi, India	1
GX	01		Guzman, France	169	JML	20		Joslin, United Kingdom	25
ICS	03		Hadhazi, Hungary	2332	JLZ	03		Juhasz, Hungary	49
IDH	03	S.	Hadhazi, Hungary	333	KMY			Kaczmarech, Brazil	2
ITY		T.	3-,	1	KPK		P.	,	2
IIVB		l.	Hajdinjak, Croatia	44 152	KB	02		Kaminski, New Mexico	4
IKB		В.	,	152	KAM	02		Kammerer, Germany	
IJW	ΩE	J.	Hall, Colorado Hambsch, Belgium	2419	KLAA		L.	Kannard, Louisiana	430
	05	F.		218130	KTU		T.	Kantola, Finland	430
HMB HHAB	02	Н.	Hammerl, Germany	55	KMO		NΛ	Kardasis, Greece	521

Table 3. AAVSO Observers, 2017–2018, cont.*

Code	Ora		Name	No. Obs.	Code	Org.		Name	No Ob.
					1	Org.			
(BJB	19		Karlsson, Sweden	809	LMA	27		Lemay, Canada	1104
THA	19	T.	Karlsson, Sweden	2418	LVY			Levy, Arizona	1.00
AD	03	Α.	1 ' 3 '	724	LPAC		P.	,	1681
EΙ		E.	Kato, Australia	153	LPIB		P.	Ligeza, Poland	17
(BJ	0.0	R.		34	LFEA	00		Limón Martínez, Spain	2
MQ	06		Kearns, Spain	3	LPEB	02	P.	Lindner, Germany	24
GAB			Kellerman, USA	806	LMK			Linnolt, Hawaii	72
HEA	20		Kerner, Germany	7	LCHF			Liponis, Greece	3
(SH	29	S.	*	10	LLEA LCO			Liponis, Greece Littlefield, Indiana	1150
CSB	03	C.	•	1852	LGIB	10		•	1150
(SZ (MR	03 20		Keszthelyi, Hungary	345	LGIB	18		Locatelli, Italy	23
HAB	20 17		Kidger, United Kingdom Kiiskinen, Finland	381 50	LOCA			Lopatynski, California	6
(THC	20		•	50 55	LVAC			Lopez, Venezuela	4
(RAA	20	T.	Killestein, United Kingdom	55 47	LCLC	10		Lopez Villanueva, France	151
(JSC		R.	5, 5	7	LRAB	18		Lopresti, Italy	151 20
(KJ	03	J.	Kirkland, Utah	4	LMTA			Loslo, Indiana	20
(KAA	03	K.	, , , , , , , , , , , , , , , , , , , ,			20		Lott, Georgia	1.7
(RAB		K.	Klindt-Jensen, Denmark	563 7	LDS LDAB	20		Loughney, United Kingdom Louw, South Africa	13
(PL		R.	1,	, 55	LFEB	02		•	1
	20	P.	Kneipp, Louisiana	448	LSJB	03 14	F.	, 3,	1
(CD	20	C.	3 ,		LBG	14		Lowther, New Zealand	226
(GT		G.	3 ,	18				Lubcke, Wisconsin	246
(SP	02	S.	3 ,	220	LLIA LPAE			Lucas, North Carolina	
(OC	03	Α.	, 3 ,	242			P.	•	
(LO		L.	•	430	LCHD			Lugova, Ukraine	53
(HL	0.2		Kohl, Switzerland	1109	LFSA		F.	Luiz, Brazil	0.7
(TAA	03	Τ.	. 3 ,	5	MDW MRGA			MacDonald, Canada	83
KMA			Komorous, Canada Konahins, Latvia	1556		02	R.	•	1.5
KALD				410	MZOA	03	Ζ.	3, ,	15
(SLA	0.2	S.	*	419	MQA		Α.	•	315
(OS	03		Kosa-Kiss, Romania	474	MDIC	02		Maitra, Massachusetts	1
(CLA	0.2	C.	· ·	1623	MII	03	L.	, , , , , , , , , , , , , , , , , , , ,	6 -
(AF	03		Kovacs, Slovakia	442	MEGA	20		Maleev, Ukraine	67
(YUB		Y.	Kozlow, Ukraine	187	MJHN	20	J.	Mallett, United Kingdom	215
KFK KJOA		F.		2	MCPA MAND			Maloney, Arkansas	187 5
	02	J.	,	88		10	Α.		
(WO	02		Kriebel, Germany	359	MXI	18		Marchini, Italy	51
(EVA	16	E.	Krotkov, Russian Federation	19	MJOE	10	J.	* •	42
NAA	02		Krumm, California	643	MFB	18	F.		
ROB	02	R.	Kubala, Germany	1	MXS	03	S.	. 3 ,	10
(BA	01	B.	· · · · · · · · · · · · · · · · · · ·	584	MMN	18		Martignoni, Italy	10
(UC	01	S.	Kuchto, France	1980		27	J.	Martin, Spain Mastellotto, Canada	
AUY		Y.	· · · · · · · · · · · · · · · · · · ·	288	MSAC	27		· · · · · · · · · · · · · · · · · · ·	
(SQ		S.	Kuznetsov, Russian Federation	29	MREN	18	R.		,
(WD			Kwadrat, Virginia	4	MISB	03	I.	•	1/
MIC	1.5		Kwieciak, Poland	69	MAV			Matsney, Russian Federation	10
.CR	15		Labordena, Spain	855	MTH	20		Matsuyama, Japan	83
HS	4.7		Lacombe, Canada	16	MDMS	20		Matthews, United Kingdom	20
SA	17		Lahtinen, Finland	104	MERA		E.		2.4
YAC			Lairgi, Morocco	53	MPR		Р.	Maurer, Germany	20
PAB	2-		Lampens-Vancauteren, Belgium	33	MZR		J.	•	3
DJ	27		Lane, Canada	1339	MCOA			McCann, Arkansas	8
TO	02		Lange, Germany	20	MTAB			McClain, Kansas	
RCA			Larochelle, Canada	5	MDP	27	P.	•	186
KR			Larsen, Connecticut	14	MCOB			McKenzie, Canada	
ZT			Lazuka, Illinois	788	MJB		J.	,	929
_JW	01		Lechopier, France	11	MMAE			McNeely, Indiana	1
_MT _PD			Legutko, Poland	155	MPJM	20	Р.	, ,	
	01	P	Lemarchand, France	3	MSD		D.	Means, Arizona	

Table 3. AAVSO Observers, 2017–2018, cont.*

Code	Ora.		Name	No. Obs.	Code	Org.		Name	No. Obs
		_			1		_		
MTOD MED			Medulka, Slovakia	8 2677	OJMA OAR	17 17	J.	, , ,	4 355
ИIQ	20	K.	,,	2677	OPR	17	P.	Oksanen, Finland	333
ліQ ЛҮАА	20	I. Y.	Megson, United Kingdom Melnikov, Austria	185 45	OSE		S.		2
MQG			Menegotto, Argentina	45 45	OJJ		э. J.		41
иQG ИZK		K.	Menzies, Massachusetts	40054	OCR	05		Otten, Belgium	40
MBEA		B.		63	PLA	05		Padilla Filho, Brazil	55
MJAL		J.	·	1	PSD		S.		93
MDEN			Merrill, California	5	PTFA		T.	, , ,	175
MHL		E.		4851	PMAJ			Papapoulias, Greece	173
MMII			Miciukiewicz, Connecticut	1	PCC	18		Papini, Italy	
MVH			Mihai, Romania	16	PPS	03	S.		1999
MIW	20	l.		18642	PSW	20		Parkes, United Kingdom	100
MMGA			Miller, Texas	64	PFRA		F.	Parks, California	114
MMEA			Millward, Australia	409	PST		S.		
MNIC			Mishevskiy, Ukraine	7772	PNIA			Paschalis, Greece	548
MOBM	20		Mobberley, United Kingdom	106	PJJ	15	J.		49
MRV		R.		17	PJRB		J.	· •	40
МНН			Moehlmann, Pennsylvania	627	PTT	20		Paterson, United Kingdom	637
MOD			Mohrbacher, Ohio	209	PGNA			Pawar, India	17
MPV	03	P.	Molnar, Hungary	2	PEX		A.	Pearce, Australia	1573
MVLA	16	V.		8	PRCA		R.		618
MISA		I.	Monks, United Kingdom	137	PMAN		M.	Peavy, California	
OAMM		M.	Morales Aimar, Spain	4729	PEI	11	E.	Pedersen, Denmark	1075
MAEA		A.	Morozov, Russian Federation	19	PVTA	11	V.	Pedersen, Denmark	17
MOW		W.	Morrison, Canada	5090	PEG	01	C.	Peguet, France	1766
MRZA		R.	Moseley, Massachusetts	3	PWD		W.	Pellerin, Texas	238
MPS	27	P.	Mozel, Canada	47	PEJA		E.	Pellett, Wisconsin	1293
MKCA		K.	Mrazek, Austria	5	PPRA		P.	Pendergraft, Alabama	76
MMH		M.	Muciek, Poland	571	PJED		J.	Penninckx, France	315
MACC		A.	Muethel, Michigan	3	PSCA		S.	Percy, United Kingdom	210
MULP	20	P.	Mulligan, United Kingdom	292	PRVA		R.	Pereira, Brazil	14
MMU		M.	Munkacsy, Rhode Island	4668	PVLA		V.	Perekhrest, Ukraine	23
MGAB		G.	•	2240	PCX	15	C.	, , ,	7
MUY	05	E.	Muyllaert, Belgium	11254	PLFA		L.		134
MGW		G.		113140	PEJ	01	J.	*	85
NDQ	01	D.	Naillon, France	32	PMAM			Perry, California	30
ATA		Α.		12	PGD			Persha, Michigan	1889
NSIA		S.	•	8	PLMA		L.	, I	3
NRNA		R.		629	PJOG		J.	Philpott, Utah	
NRRB			Neibaur, North Carolina	2	PTEB			Phong, Australia	264
NLX		P.	Nelson, Australia	14845	PXR	20	R.	Pickard, United Kingdom	15922
NTHC		T.		41	PKT		J.	· ·	364
NAL	02	Α.	, 3 ,	421	PROC		R.	'	216
VJO	02	J.		2703	PUWA			Pilz, Germany	10
IMI	02		Nicholas, Arizona	2142	PRUA			Pinizzotto, Maine	37
TON	02		Nickel, Germany	1952	PJEF	02		Pioppa, France	364
TMV	17		Nissinen, Finland	130	PIJ PMAA	03	J.	. 3 ,	33
AMON			Noroozi, Michigan	396				Pirtac, Romania	(
NCH			Norris, Texas Novichonok, Russian Federation	61 65	PMAF PWMA			Pittendreigh, Florida Pittendreigh, Florida	16
NAO DCN		A. S.	,	65 292	PVVIVIA		vv. P.	3 ,	160
			O'Keeffe, GA			20			132 2813
DBRA			,	3 97	PAW	29 12		Plummer, Australia	
odea Onj		D. J.	O'Keeffe, Ireland O'Neill, Massachusetts	87 709	AST PTOB	12	R.	Podesta, Argentina Polakis, Arizona	3 290
OJEA		J.		798 319	PVEA		۱. V.		1469
			Odasso, Italy				v. J.		140
$\gamma \Lambda \varsigma$		Α.	Ouasso, italy	1	PJOF		J.	r oppeie, minnesota	1.
oas Oela		E.	Odintsova, Ukraine	180	PJTA	17	J.	Porio, Finland	8

Table 3. AAVSO Observers, 2017–2018, cont.*

C04-	0		Nama	No.	Codo	0:		Namo	No.
Code	Org.		Name	Obs.	Code	Org.		Name	Obs
POX		M.	Poxon, United Kingdom	1024	SIRC		I.	Sagaybins'ka, Ukraine	209
PYG	20		Poyner, United Kingdom	10635	SSU		S.	Sakuma, Japan	21
PEMB			Primucci, Argentina	431	SFV	18	F.	33 , ,	6
PGB			Profita, Italy	8	SBAH		В.	Salwiczek, Poland	31
PALA			Prokofyev, Cyprus	4	SAH			Samolyk, Wisconsin	9552
PMB			Prokosch, Texas	3	DSS	06	Α.	3 3 .	
PDQ	01		Proust, France	97	SFGA		F.		1
PJUA		J.	Provencal, Delaware	2	SSIM		S.		7
PUJ	06	F.	Pujol-Clapes, Spain	665	SGE	27		Sarty, Canada	
PARA		Α.	27 1	30	SJOG		J.	Saunders, New York	50
PHG			Purucker, Germany	2	SEDB		Ε.	, ,	50
PALE		Α.		1254	SMAI			Sblewski, Germany	24.
PMAK	10		Pyatnytskyy, Ukraine	1577	SDAV			Scanlan, United Kingdom	41
QULA	18		Quadri, Italy	6850	SFS SRBR		S.	, 3	719
QCHA	02		Quesada, Arizona	2	SPK	01		Schippers, Netherlands	330
RKE	02	K.	Raetz, Germany	541	SHV	03	P.	Schmeer, Germany Schmidt, Hungary	1
RJOC		J.	Rallo, Spain	42 69	SREB	03		Schmidt, District of Columbia	303
RBK RGAB		B.	Ramotowski, New Mexico Ramsay, United Kingdom	1	SRAB	02		Schoenfeld, Germany	16
RMAF		G. M.	,	559	SFRA	02		Schorr, Georgia	196
RACA		Α.		9	SYU	02		Schubert, Germany	133
RRUB		R.	Rasmussen, Denmark	4	SBEA	02	B.		36
RRIB	27	R.		4	SDM	02		Schwendeman, Virginia	81
RMN	21		Ratcliffe, Kansas	3	SJEA	01	J.	. 2	37
RJEA		J.	Rayon, France	2	SJTS	20	J.		4247
RWSA	29		Rea, New Zealand	1167	SJIA	20	J.		1171
REP	24	P.	Reinhard, Austria	296	SHUA			Sears, Ohio	61:
RFP	13	P.	Reis Fernandes, Brazil	181	SJPA		J.	Seitz, Michigan	209
RJG	13	J.	Ribeiro, Portugal	1129	SSAB		S.		207
RIX	29	T.	Richards, Australia	560	SASC		Α.		12
RHM			Richmond, New York	33302	SJOI		J.	•	9
RCCA		C.	Riou, France	89	SIV		I.	Sergey, Belarus	49
DJR		J.	Ripero Osorio, Spain	3713	SFL		F.	Sevilla Lobato, Spain	79.
RIZ		J.	Ritzel, New York	4160	SSTA	27	S.	Shadick, Canada	26
RNDA		N.		345	SSHA		S.		240
RJJC		J.	Rivet, Texas	58	SJDA	20	J.	Shanklin, United Kingdom	47:
RRIA		R.	Roberts, Virginia	687	SHS		S.	Sharpe, Canada	269
RJWA		J.	Robertson, Arkansas	364	SQN		L.	Shaw, California	14
RPT		P.	Rochford, Alabama	10	SFY	20	J.	Shears, United Kingdom	4460
RJWB		J.	Rock, United Kingdom	29297	SVLA		V.	Shlyonskov, Russian Federation	
RAEA		A.	Rodda, United Kingdom	314	SLH		L.	Shotter, Pennsylvania	46
BRB		B.	Rodgers, Canada	36	SGQ		C.	Sigismondi, Italy	8
RHE	26	Н.	Rodriguez, Uruguay	5	SPAO	18	P.	Siliprandi, Italy	46
ZD		D.	Rodriguez Perez, Spain	318	SCMA		C.	Silva, Argentina	
RDAE		D.	Rodriguez Vides, Spain	55	SBN	13	A.	Silva Barros, Brazil	
rjad		J.	Rodríguez, Spain	5	SGEO			Silvis, Massachusetts	171
RANC		A.	Roerig, Germany	1	SNE			Simmons, Wisconsin	290
RRO		R.	,	1	SMLA			Simonson, Washington	3
RES		E.	•	218	SANG			Sing, Philippines	71
RACB		A.	•	1	SGOR			Sjöberg, Massachusetts	594
RIC		R.	•	6	SAJD			Skoczypiec, Poland	1
OG			Ross, Michigan	147	SDN			Slauson, Iowa	15
RGN			Rossi, Italy	252	SDAB			Smales, United Kingdom	21
JV		J.	Ruiz Fernandez, Spain	990	STAC			Smela, Poland	9
RMTB			Russell, Colorado	130	SDWA	20		Smith, United Kingdom	233
RUJ		J.	Ruthroff, Indiana	1072	SDZ			Smith, Arizona	1
RRR		R.	, ,	32	SHA			Smith, Michigan	54
2711		M.	Rzepka, Poland	2	SVIC			Smolianchuk, Ukraine	2
RZM SRIC		_	Sabo, Montana	32253	SKMA			Sniegocki, Poland	

Table 3. AAVSO Observers, 2017–2018, cont.*

				No.					No.
Code	Org.		Name	Obs.	Code	Org.		Name	Obs
SLUA		L.	Socha, Poland	139	TVAB		V.	Tsehmeystrenko, Ukraine	74
SROD		R.	Solomon, Australia	625	TANB			Tsvetkov, Russian Federation	2
YRA		Y.		164	TYEA		Y.		17
TFB		T.	. '	4	TGOB			Tucker, California	
OI			Soukup, New Mexico	86	TGR	27		Tulloch, Canada	84
TYA		Т.		23	TADA			Tużnik, Poland	
ilQ			Spearman, Texas	15	TYS			Tyson, New York	18
PGA		P.	•	7	TSAA			Tzikas, Virginia	68
SXR	03		Sragner, Hungary	7	UAN	03	Α.		208
BL	05		Staels, Belgium	2261	UJHA	05	J.	Ulowetz, Illinois	4861
VAE		٧.		1	UIS01			Barber Obs. U. Illinois, Illinois	94
		_	. 3						
DB		D.	· · · · · · · · · · · · · · · · · · ·	8247	UMAA	01		Urbanik, Slovakia	518
PET		P.		10083	VLN	01		Vadrot, France	
ABB		Α.	• *	104	VSEA			Vakulenko, Ukraine	15
NIB		N.		1	VRUC		R.	•	_
ET			Stephan, West Virginia	142	SROF			Valdomiro Silva, Brazil	6
CEA			Stewart, Washington	161	VTY	20		Vale, United Kingdom	81
CMB		C.	Stinson, Virginia	1	VADA		A.	Valvasori, Italy	
OX		C.	Stockdale, Australia	678	BVE	04	E.	Van Ballegoij, Netherlands	130
GEA		G.	Stone, California	220230	VBRB		B.	Van Deventer, Washington	(
MIF		M.	Stone, Missouri	153	VDE	04	E.	Van Dijk, Germany	11
DI	20	D.	Storey, United Kingdom	82	VWS	05	J.	Van Wassenhove, Belgium	373
IAK		I.		3457	VBH	05	Н.	Vandenbruaene, Belgium	
HZ	02	Н.	Struever, Germany	27	VMAE	05		Vanleenhove, Belgium	120
RX	14	R.	Stubbings, Australia	22	VMT	05		Vanmunster, Belgium	11692
AC	02		Sturm, Germany	858	VSD	05		Vansteelant, Belgium	
SUS	02		Suessmann, Germany	685	VANA			Vasilev, Ukraine	884
SPP		P.		79	VED	01	Р.	•	4945
SJAR		J.	Suomela, Finland	177	VRG	01		Venne, Canada	10
SB		R.		1	VBPA			Vietje, Vermont	1147
SWV			Swann, Texas	403	VII	03	I.	Vincze, Hungary	4
TDA		T.		2	VGK	05		Vithoulkas, Greece	100
	02			48	VGK	02			
SZX	03		Szalma, Hungary			02	F.		977
AO	03		Szauer, Hungary	118	VOL			Vollmann, Austria	224
SLY	03	L.	Szegedi, Hungary	119	VMAG			Vrastak, Slovakia	10
TFUA		F.		50	WEO		E.	3 .	-
TUO .		U.	3 · ,	59	WWRA			Waddell, North Carolina	
MAA			Talero, Spain	56	WCR	27		Wagner, Canada	11:
CGA	20	C.	, ,	27	WGR			Walker, New Hampshire	267
DB	27		Taylor, Canada	180	WIV	20	I.	Walton, United Kingdom	12
SUB		S.	Taylor, Colorado	16	WGAB		G.	Wandless, Virginia	
ΓLMA		L.	Teixeira, Brazil	11	WJOB	19	J.	Warell, Sweden	3
PS	03	I.	Tepliczky, Hungary	1230	WAU		A.	Wargin, Poland	
DN		D.	Terpstra, Arizona	132	WTIB		T.	Weber, Colorado	
TU		T.	Tezel, Turkey	4	WPT		P.	Wedepohl, South Africa	2
JP	20	J.	Thorpe, Australia	68	WRCA		R.		7.
JD		J.		137	WWC			Weiss, Arizona	
IΑ	03	A.	_	260	WLIA		L.		
SCB	0.5	S.	, 3 ,	78	WKL	02	K.	•	91
RE		R.		6453	WROC	02		Werder, Germany	23
NIA		N.		215	WJAA		J.		26
00	20		Toone, United Kingdom			20		White, United Kingdom	
	20	J.		6458	WNIB	20			20
ST	02	S.	•	53026	WBOA	20		Wiskens United Kingdom	28
RT	03	T.	, , ,	53026	WNEA	20		Wickens, United Kingdom	7.
VT			Tramazzo, Arizona	4	WWK			Wierzchos, Florida	18
RF			Trefzger, Switzerland	137	WTHB	19		Wikander, Sweden	219
RH	20	R.	· · · · · · · · · · · · · · · · · · ·	124	WEY			Wiley, Texas	3806
DW		D.	Trowbridge, Washington	14	WTEB		T.	Willamo, Finland	7
			Tsao, Taiwan					Williams, Ohio	

Table 3. AAVSO Observers, 2017-2018, cont.*

Code	Org.	Name	No. Obs.	Code	Org.	Name	No. Obs.
WPX	29	P. Williams, Australia	6601	ZGEA		G. Zhao, California	5
WSN		T. Wilson, West Virginia	1319	ZANA		A. Zhornichenko, Russian Federation	1
WERB	02	E. Wischnewski, Germany	458	ZUD		D. Zubenel, Kansas	140
WKM		M. Wiskirken, Washington	5				
WPB	20	P. Withers, United Kingdom	1573				
WGI	02	G. Wollenhaupt, Germany	6				
WTOC		T. Wrocławski, Poland	1				
WUB	04	E. Wubbena, Netherlands	410				
WCG		C. Wyatt, Australia	78				
YIGA		I. Yatsenkov, Russian Federation	231				
YBA		B. Young, Oklahoma	321				
YCRA		C. Young, New Zealand	8198				
YDV		D. Young, Massachusetts	3000				
YON		R. Young, Pennsylvania	16				
ZMOA		M. Zachariasen, Denmark	5				
ZALB	37	A. Zanardo, Brazil	2				
ZPA		P. Zeller, Indiana	2				

^{*}Totals reflect observations made during fiscal 2017–2018 and do not include historical data (data preceding fiscal 2017–2018) submitted during fiscal 2017–2018.

These codes, which appear in the Table (AAVSO Observers 2017–2018), indicate observers are also affiliated with the groups below:

- 01 Association Française des Observateurs d'Étoiles Variables (AFOEV)
- 02 Bundesdeutsche Arbeitsgemeinschaft für Veränderliche Sterne e.V. (BAV) (Germany)
- Magyar Csillagàszati Egyesület, Valtózocsillag Szakcsoport (Hungary)
- 04 Koninklijke Nederlandse Vereniging Voor Weer-en Sterrenkunde, Werkgroep Veranderlijke Sterren (Netherlands)
- 05 Vereniging Voor Sterrenkunde, Werkgroep Veranderlijke Sterren (Belgium)
- 06 Madrid Astronomical Association M1 (Spain)
- 08 Norwegian Astronomical Society, Variable Star Section
- 10 Astronomical Society of Southern Africa, Variable Star Section
- 11 Astronomisk Selskab (Scandinavia)
- 12 Liga Iberoamericana de Astronomia (South America)
- 13 Rede de Astronomia Observacional (Brazil)
- $14\quad Royal\,Astronomical\,Society\,of\,New\,Zealand, Variable\,Star\,Section$

- 15 Agrupacion Astronomica de Sabadell (Spain)
- 16 Association of Variable Star Observers "Pleione" (Russia)
- 17 URSA Astronomical Association, Variable Star Section (Finland)
- 18 Unione Astrofili Italiani (Italy)
- 19 Svensk Amator Astronomisk Förening, Variabelsektionen (Sweden)
- 20 British Astronomical Association, Variable Star Section
- 24 Astronomischer Jugendclub (Austria)
- 26 Red de Observadores (Montevideo, Uruguay)
- 27 Royal Astronomical Society of Canada
- 29 Variable Stars South (New Zealand)
- 31 Center for Backyard Astronomy
- 32 Groupe Europeen d'Observations Stellaires
- 34 Astronomical Society of South Australia
- 36 Nucleo de Estudo e Observacao Astronomica--Jose Bazilicio de Souza (Florianopolis, Brazil)
- 37 Clube De Astronomia De Sao Paolo (Brazil)



Appendix 3. Volunteers

AAVSO Volunteers

AAVSO members are very generous with their time and talents. Many of the programs and services we offer would not be possible without the participation of member volunteers: they are regularly involved in teaching new observers; writing articles for our publications; helping to keep the *Variable Star Index* up to date and functional, and submissions vetted; and the creation of charts and comparison star sequences.

We take this opportunity to recognize these special people who volunteered during the fiscal year, and to say *thank you* for another year of valuable contributions of time and expertise.

Mentor Program

Patrick Abbott	Ron Fournier	Ken Menzies
Umair Asim	Jim Fox	Gordon Myers
	Bill Goff	Peter Nelson
David Benn	Keith Graham	Sebastián Otero
John A. Blackwell	Barbara Harris	Roger Pieri
John Centala	Stella Kafka	Brad Vietje
Michael Cook	Roger Kolman	William Wilson
Michael Deconinck	Mike Linnolt	vviiidili vvii3011

Variable Star Index (VSX) Team

Charts and Sequences

Tom Bretl	Arne Henden	Matthew Templeton
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Jack Davis	Mati Morel	Natalia Virnina
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Observing Manuals

AAVSO Solar Observing Guide - Version 1.1

Frank Dempsey Kristine Larsen
Raffaello Braga Ralph Chou
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Translators

Photoelectric Photometry Observer's Guide French - Bernard Candela

A Practical Guide to Exoplanet Observing French - Jean-Bruno Desrosiers, Serge Bergeron, and Manon Bouchard

Solar Observing Guide

French - Bernard Candela German - Ilka Petermann Italian - Raffaello Braga Portuguese - Alexandre Amorim Spanish - Sebastián Otero

Variables: What are they and why observe them? Uzbek - Sherali Niyazova



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Friedrich Wilhelm August Argelander (1799–1875)

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